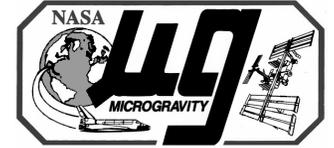




Working in a Reduced Gravity Environment : “A Primer”



SECTION 4

Working in a Reduced Gravity Environment: “A Primer”

Kenol Jules

PIMS Project Scientist

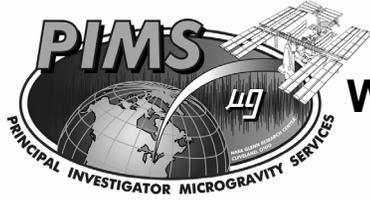
NASA Glenn Research Center

March 4th, 2003



CONTENT

- ▶ **Reduced Gravity Environment Description
--- Overview**
- ▶ **ISS Microgravity Environment Requirements**
- ▶ **Microgravity Disciplines Sensitivity Assessment**
- ▶ **Vibration Isolation--- Why?**
- ▶ **ISS Coordinates Systems**
- ▶ **Plots Header Description: ISS**
- ▶ **Experiment Planning and Execution**
- ▶ **Preflight Planning for Science Optimization**
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- ▶ **Plots Header Description: ISS**
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INTRODUCTION

- **Earth-bound experiments are affected by normal gravity and vibrational forces, which exist in ground laboratories:**
 - ▣ **gravity, elevators, air conditioner, people and so on...**
- **Most microgravity experiments desire:**
 - ▣ **zero gravity, or**
 - ▣ **constant, uni-directional acceleration, or**
 - ▣ **constant conditions**
- **Taking experiments to orbit removes effects of gravity, but trades ground disturbances for other:**
 - ▣ **gravity gradient, aerodynamic drag, thrusters, other experiments disturbances, vehicle sub-systems and crew disturbances**
- **Experiments may be disturbed by accelerations from various sources**



REDUCED GRAVITY ENVIRONMENT DESCRIPTION

The reduced gravity acceleration environment of an orbiting spacecraft in a low earth orbit is a very complex phenomenon. Many factors contribute to form the overall environment. In general, it can be considered as made up of the following three components:

QUASI-STEADY: is composed of those accelerations that vary over long periods of time, typically longer than a minute for space-based platforms.

VIBRATORY: is composed of those accelerations that are harmonic and periodic in nature with a characteristic frequency.

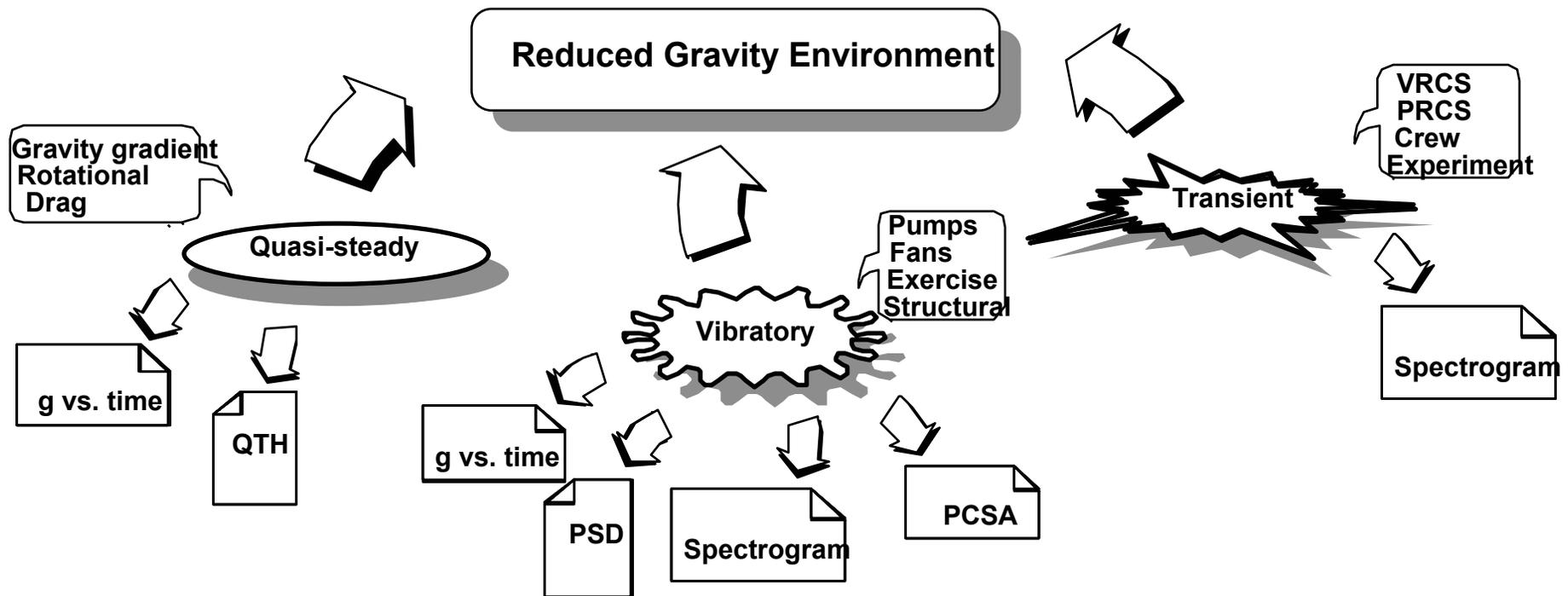
TRANSIENT: is composed of those accelerations that last for a short period of time, and are non-repetitive.

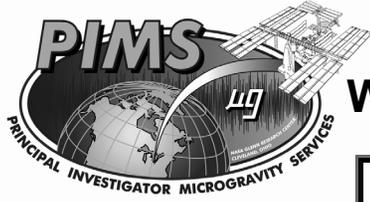
REDUCED GRAVITY ENVIRONMENT DESCRIPTION

What is a "Reduced Gravity Environment" ?

Major properties

- What causes these accelerations
- ▭ How do we display them?





Working in a Reduced Gravity Environment : “A Primer”

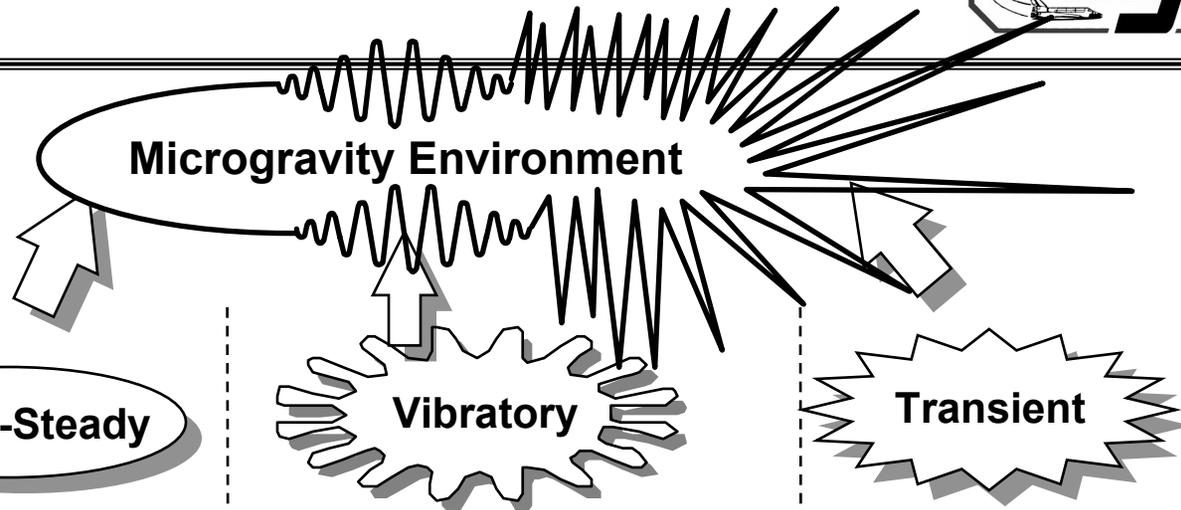


REDUCED GRAVITY ENVIRONMENT DESCRIPTION

Definitions

- **Reduced Gravity Environment:** an environment in which the effects of gravity are small compared to those we experience on Earth
- **Quasi-steady:** signal which varies at a very low frequency, typically below 0.01 Hz
- **Vibratory:** term used to describe vibratory disturbances with frequency content greater than 0.01 Hz
- **Transient:** signals that are impulsive in nature; passing quickly into and out of existence. They are broadband.

Microgravity - generic term applied to scientific investigations that exploit or explore near weightlessness; not a reference to a specific acceleration level



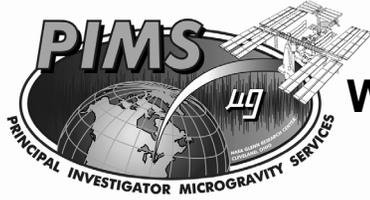
Component	Quasi-Steady	Vibratory	Transient
Instrument	MAMS	SAMS & MAMS	SAMS & MAMS
Frequency	$0 \leq f < 0.01$ Hz	$0.01 \leq f \leq 300$ Hz	broadband
Magnitude	μg 's (or less) peak	tens to thousands of μg_{RMS}	tens of mg's peak
Primary Sources	<p>gravity gradient & rotational effects: not at center of mass</p> <p>drag: function of altitude, attitude, day/night, etc.</p> <p>vehicle: venting water or air</p>	<p>equipment: pumps, fans, centrifuges, compressors, etc.</p> <p>crew: ergometer or treadmill exercise</p> <p>vehicle: structural modes</p>	<p>vehicle: thrusters, dockings</p> <p>crew: pushoffs & landings, drawer/door closings, experiment setup</p> <p>equipment: machinery startup</p>
Microgravity Requirement	<p>Magnitude - $1.0 \mu\text{g}$</p> <p>($0 \leq f \leq 0.01$ Hz.)</p> <p>Stability - $0.2 \mu\text{g}$ perpendicular component to orbital average QS acceleration vector</p>	<p>Combined Vibratory - per figure ($0.01 \leq f \leq 300.0$ Hz.)</p> <p>•100 second root mean square average</p> <p>•Per one-third octave band</p>	<p>Individual Transient</p> <p>•1000 μg peak per axis</p> <p>•10 $\mu\text{g}^*\text{s}$ integrated over any 10 s interval per axis</p>



REDUCED GRAVITY ENVIRONMENT DESCRIPTION

WHAT DO ALL THESE MEAN TO YOU?

- The environment is NOT “zero-g”!
- Experiments may be affected by the reduced gravity environment
- This tutorial will explain to you what the environment is likely to be, how we measure it, how we interpret it, and will show you what impact the environment has had on some experiments.



CONTENT

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--- Overview**
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- ▶ **ISS Coordinates Systems**
- ▶ **Plots Header Description: ISS**
- ▶ **Experiment Planning and Execution**
- ▶ **Preflight Planning for Science Optimization**
- ▶ **Overall Summary**
- ▶ **Extra Information: STS Ascent and Landing Profiles**



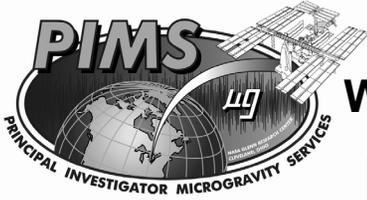
ISS Microgravity Environment Requirements

- **Microgravity science researches perform on the ISS are science “requirement” driven, rather than vehicle “capability” driven**
- **Space Shuttle is used for microgravity experiments, but there are no Space Shuttle microgravity requirements**
- **ISS, on the other hand, DOES have microgravity requirement associated with the different regimes: Quasi-steady, Vibratory, Transient and time duration. These requirements are discussed in the next few charts**



ISS Microgravity Environment Requirements

- The environmental requirement is specified as a “not to exceed” acceleration magnitude
 - Quasi-steady accelerations have directional requirements
 - Vibratory accelerations have RMS limits as a function of frequency
 - Transient accelerations have both magnitude limitation and integrated acceleration requirements
- These requirements must be met in minimum time intervals of 30 continuous days, with a cumulative time duration of not less than 180 days per year
- These must be achieved at 50% or more at the designated internal science locations (racks) on the station



ISS Microgravity Environment Requirements

- During the periods designated as “Microgravity mode”, ISS will be operated to meet the microgravity environment requirements. Otherwise, ISS will be in what is called: “Non-microgravity mode”
 - These requirements applied only for ISS assembly complete
- Design requirements and operational constraints on ISS are intended to limit acceleration disturbances in three regimes:
 - Quasi-steady, $f \leq 0.01$ Hz
 - Vibratory, $0.01 \text{ Hz} \leq f \leq 300$ Hz
 - Transient (short duration relative to an orbital period, non-periodic and broadband)



Quasi-steady

- **Frequency content: DC to 0.01 Hz**
- **Three main components of QS Vector**
 - **Aerodynamic Drag (Spacecraft wetted surface area)**
 - Attitude
 - Atmospheric density (time and altitude dependent)
 - ISS Configuration
 - **Rotational Effects (Spacecraft rotating about its center of mass-CM)**
 - Attitude
 - Angular velocity
 - Position relative to ISS Center of Mass
 - **Gravity Gradient (Experiment not necessarily located at the CM)**
 - Attitude
 - Position relative to ISS Center of Mass
- **Disturbances in the Quasi-Steady Environment**
 - **Crew Activity Effects**
 - **Air and Water Venting**
 - **EVA/SSRMBS Operations**
 - **Miscellaneous**



Working in a Reduced Gravity Environment : “A Primer”



Vibratory

- **Crew**

- Sleep/Wake
- Exercise

- **Vehicle**

- Docking
- Air conditioner [SKV]; Система Кондиционирования Воздуха (СКВ)
- Vehicle structural modes
- Etc ...

- **Experiment**

- Experiment of Physics of Colloids in Space (EXPPCS)
- ADVanced ASTroCulture (ADVASC)
- Gas Analysis System for Metabolic Analysis of Physiology (GASMAP)
- Microencapsulation Electrostatic Processing System (MEPS)
- Characterization Experiment (ARIS-ICE) Hammer Test
- Active Rack Isolation System International Space Station (ARIS-ICE)
- Etc ...



ISS Microgravity Environment Requirements

ISS Microgravity Requirements

Summary

Quasi-steady

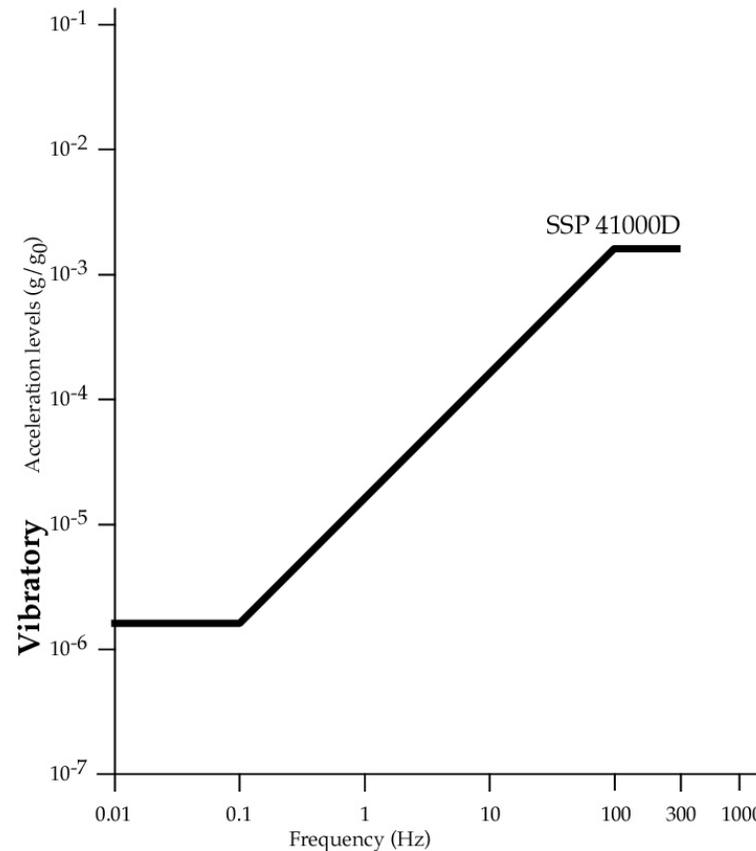
- Steady state $f < 0.01$ Hz
- $g \leq 1 \mu g_{rms}$
- Stability: perpendicular $g \leq 0.2 \mu g_{rms}$

Vibratory

- Levels in figure at structural mounting interfaces
- RMS acceleration magnitude in one-third octave averaged over 100 seconds
- Does not include crew disturbances

Transient

- $g \leq 1000 \mu g$ per axis
- $g \leq 10 \mu g\text{-sec}$ per axis (integrated over 10 sec)



For $0.01 \leq f \leq 0.1$ Hz : $a \leq 1.6^* \mu g$ or $1.8^{\$} \mu g$
 For $0.1 < f \leq 100$ Hz : $a \leq f \times 16 \mu g$ or $18 \mu g$
 For $100 < f \leq 300$ Hz : $a \leq 1600 \mu g$ or $1800 \mu g$
 Where: f = frequency
 a = acceleration

* Vehicle alone
 \$ Vehicle + payload

March 4th, 2003

ISS Microgravity Environment Requirements

ISS Microgravity Environment

THE Requirement for the International Space Station

DURATION

Mode: Microgravity - habitable

This mode consists of capabilities required for microgravity research by user payloads in a habitable environment. This mode does not include the effects of crew activity, but does include the effects of crew equipment, such as the operation of exercise devices and latched or hinged enclosures. Crew effects will be mitigated to the extent possible. This mode consists of the capabilities described in SSP 41000 and the following unique capability.

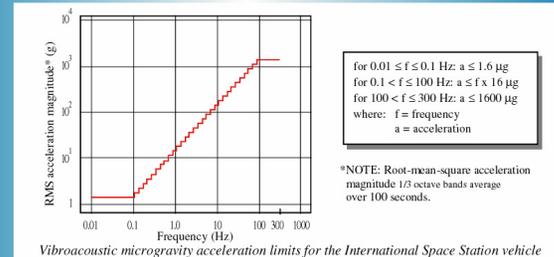
Capability: Support microgravity experiments

The purpose of this capability is to establish the required environment for microgravity experiments. The Space Station shall provide the following microgravity acceleration performance for at least 50 percent of the internal payload locations (excluding Nadir window payload location) for 180 days per year in continuous time intervals of at least 30 days:

- a. At the centers of the internal payload locations, a quasi-steady (<0.01 Hz) acceleration
 - (1) Magnitude less than or equal to 1 micro-g
 - (2) Component perpendicular to the orbital average acceleration vector less than or equal to 0.2 micro-g
- b. At the structural mounting interfaces to the internal payload locations
 - (1) A vibratory acceleration limit as defined in the figure below
 - (2) A transient acceleration limit for individual transient disturbance sources less than or equal to 1000 micro-g per axis
 - (3) An integrated transient acceleration limit for individual transient disturbance sources less than or equal to 10 micro-g seconds per axis over any 10 second interval

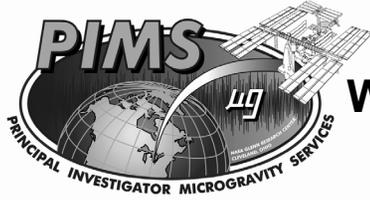
The Space Station shall monitor and record the microgravity environment at selected locations.

VIBRATORY



QUASI-STEADY

TRANSIENT



CONTENT

- ▶ **Reduced Gravity Environment Description
--- Overview**
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- ▶ **Microgravity Disciplines Sensitivity Assessment**
- ▶ **Vibration Isolation--- Why?**
- ▶ **ISS Coordinates Systems**
- ▶ **Plots Header Description: ISS**
- ▶ **Experiment Planning and Execution**
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EXPERIMENT SENSITIVITY ASSESSMENT

Fundamental Physics

Quasi-steady

- A large quasi-steady level will destroy sample uniformity of critical fluid

Vibratory

- Primary concern is vibratory heating of sample and destruction of sample uniformity

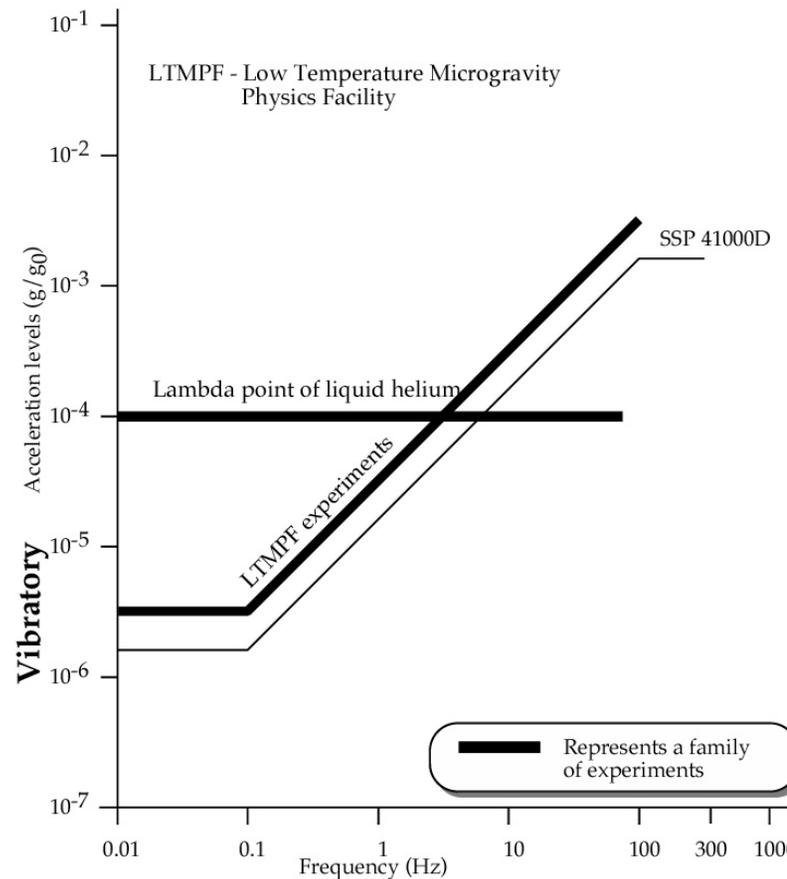
Transient

- Primary concern is vibratory heating of sample and destruction of sample uniformity

Rationale

- Low temperature physics experiments rely on establishment of highly uniform sample in microgravity

-
- NOTE: Many of these experiments are expected to be operated on the JEM-EF



EXPERIMENT SENSITIVITY ASSESSMENT

Combustion Science

Quasi-steady

- Not a major concern ($10^{-4} g_0$)

Vibratory

- Typically low acceleration levels at low frequencies (< 1 Hz) disturb experiments
- Most experiments are above the ISS requirement curve but some are below the expected environment
- Low frequency g-jitter has been observed repeatedly to affect the combustion characteristics of a variety of flames, e.g., candle, gas jet, flame balls, etc.

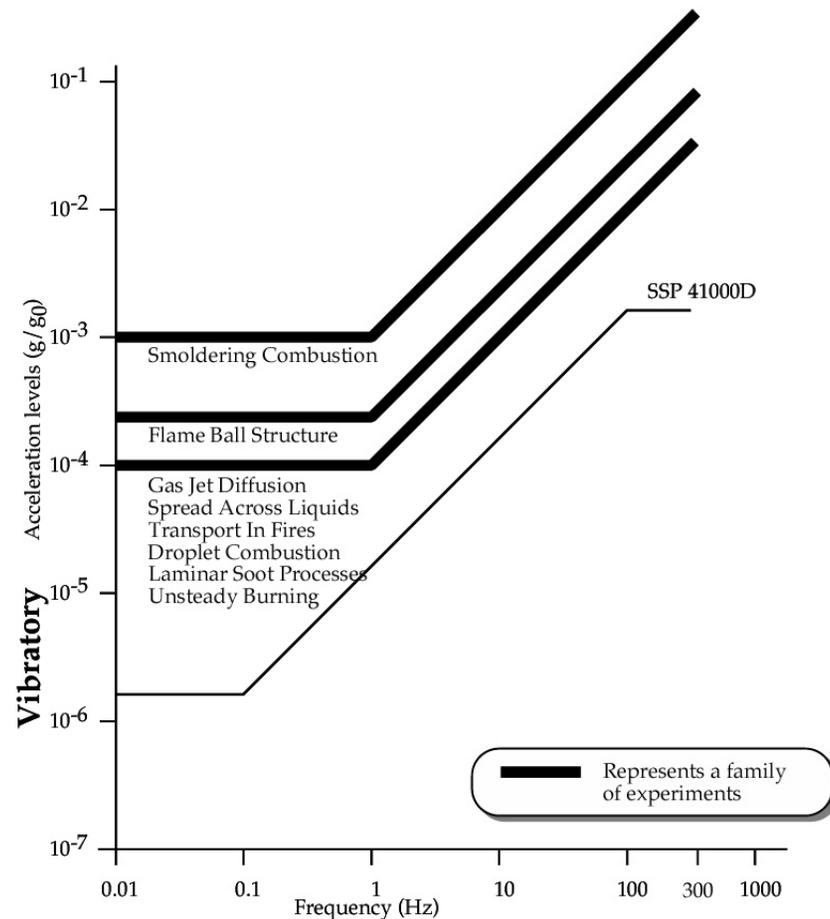
ref: Dr. H. Ross/NASA LeRC

Transient

- Tolerable for most experiments with time and magnitude restrictions on the disturbance

Rationale

- Microgravity conditions allow:
 - isolation of gravity-driven mechanisms;
 - influence of transport phenomena
 - creation of symmetry and/or boundary & initial conditions
 - new diagnostic probing or testing of similitude
- Microgravity environment has attracted widespread external peer advocacy for combustion science in space





EXPERIMENT SENSITIVITY ASSESSMENT

Biotechnology

Quasi-steady

- Not a major concern (10^{-3} to 10^{-4} g_0)

Vibratory

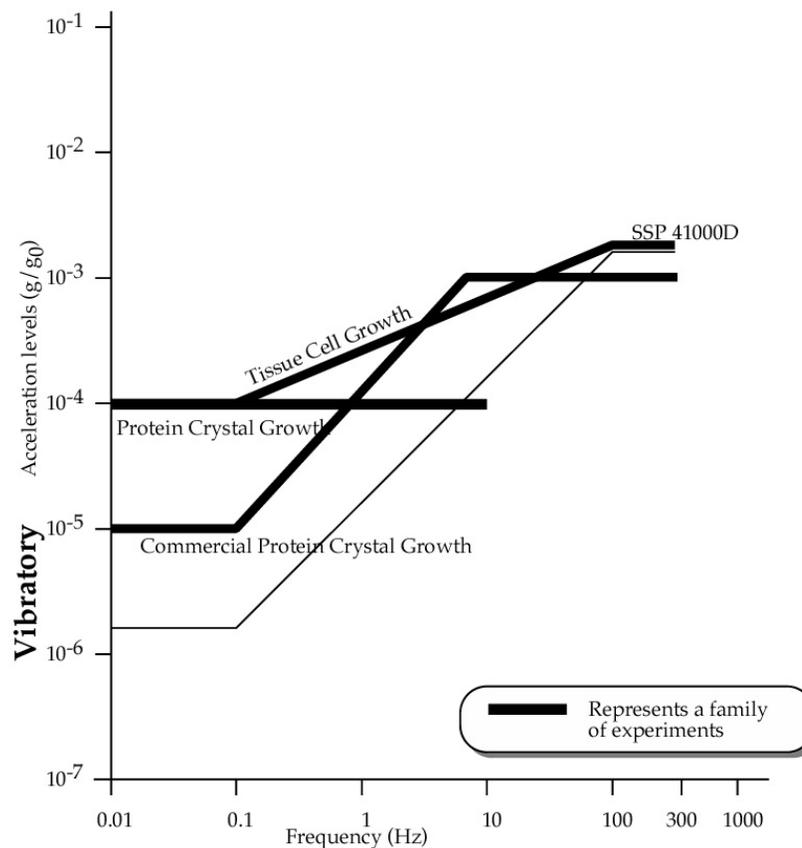
- Impact at higher frequencies of the desired operating level

Transient

- Primary concern is for large scale accelerations, such as Orbital Maneuvering System engines and crew disturbances

Rationale

- Large disturbances cause nucleations to occur in multiple sites destroying single crystal formation



EXPERIMENT SENSITIVITY ASSESSMENT

Fluid Physics

Quasi-steady

- Quasi-steady accelerations disturb most fluid experiments ($2 \times 10^{-6} g_0$)

Vibratory

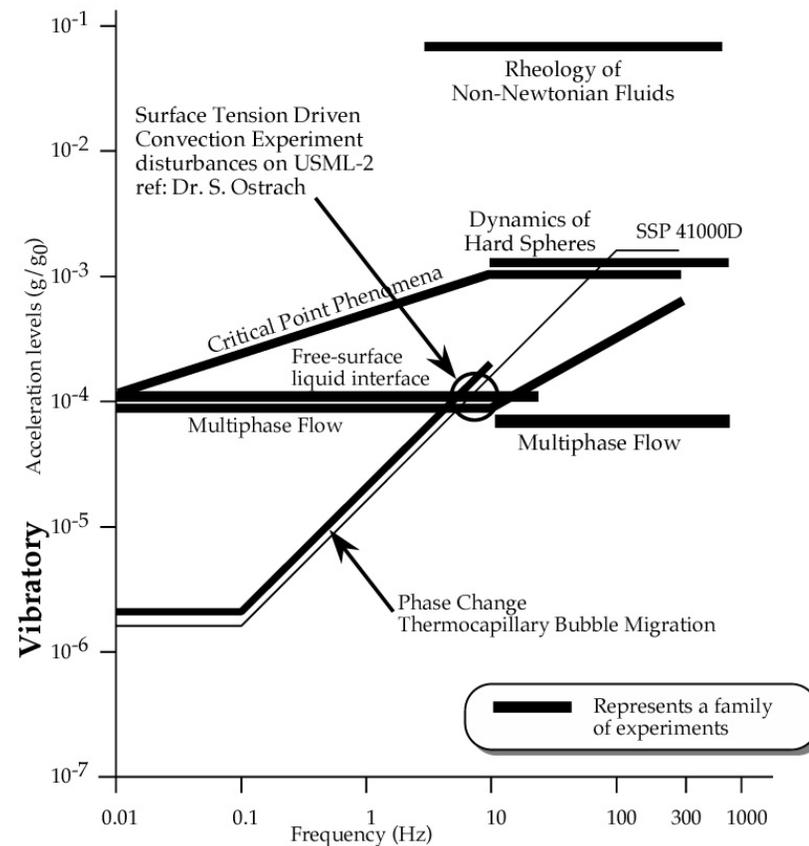
- Mid-range frequencies of expected environment will disturb fluid free surface experiments
- Some experiments require environment at lower levels than the ISS requirements curve e.g. Thin Film Fluid Flows at Menisci
- Surface Tension Driven Convection Experiment experienced surface distortion due to g-jitter frequently throughout the USML-2 mission
ref: Dr. S. Ostrach/CWRU

Transient

- Transients disturb fluid experiments with lower viscosity fluids

Rationale

- Accelerations above a threshold cause interface instability, density settling, and density-driven convection & mixing



EXPERIMENT SENSITIVITY ASSESSMENT

Materials Science

Quasi-steady

- Some samples and processes require very low quasi-steady acceleration levels ($a < 0.1$ micro-g)
e.g., Stoke's settling, Bridgman growth, Float zone
- Residual acceleration direction and stability are important factors for crystallization processes
- A Crystal Growth Furnace sample was withdrawn from USML-2 due to a change in Orbiter attitude just before launch
ref: Dr. S. Lehoczky/NASA MSFC

Vibratory

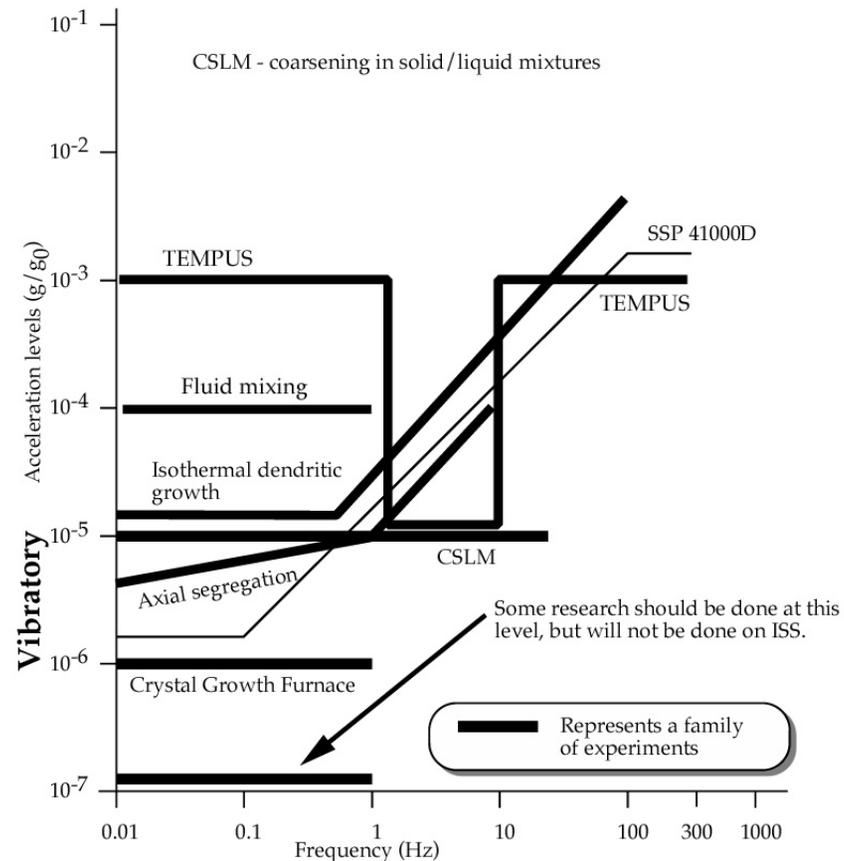
- Disturbances in various frequency ranges disturb experiments involving molten samples, suspended samples, etc.

Transient

- Some processes are very susceptible to transients such as thruster firings
- MEPHISTO (USMP-1 & USMP-3) experienced effects which lasted minutes from single thruster firings (0.01 g for 10 - 25 seconds)

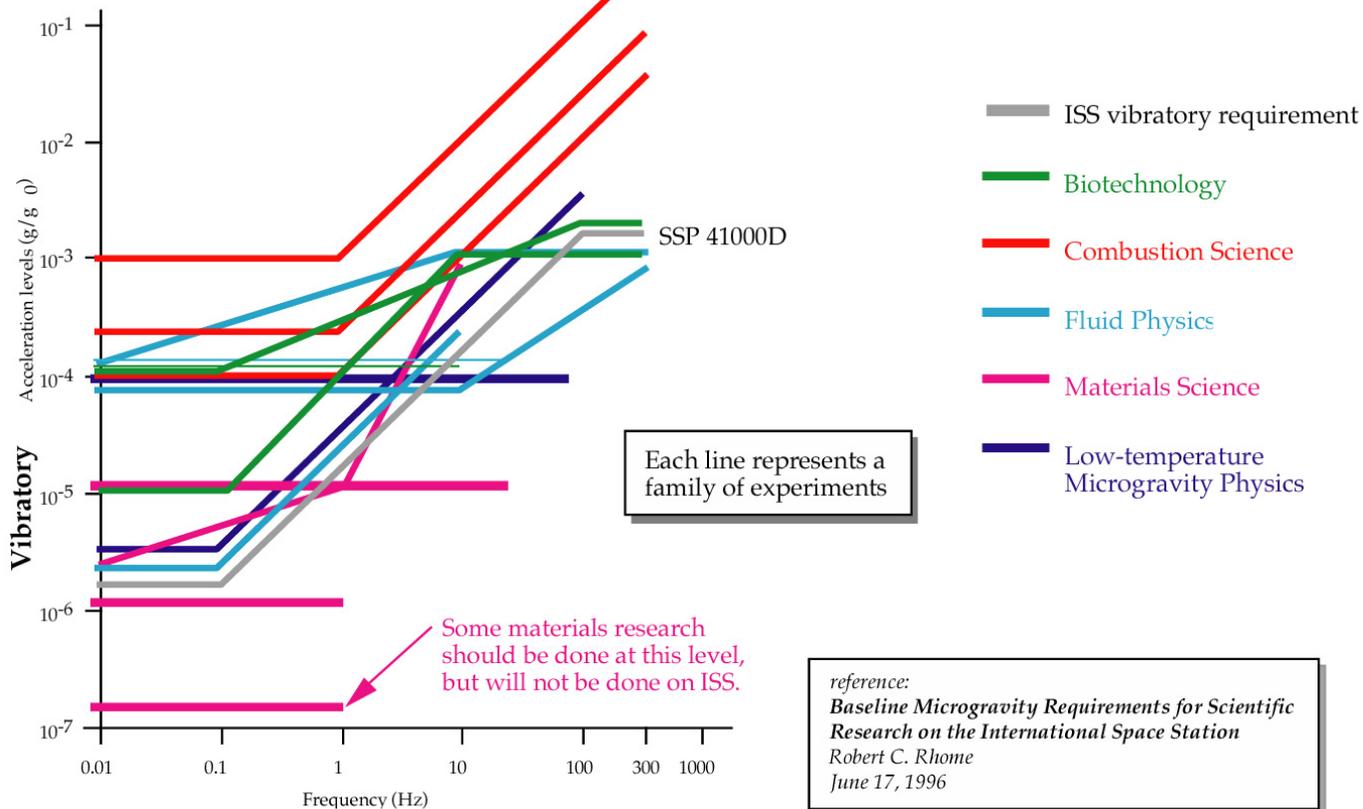
Rationale

- Accelerations above a threshold cause thermo-solutal convection and interface instability



EXPERIMENT SENSITIVITY ASSESSMENT

Science Discipline Microgravity Vibration Requirements

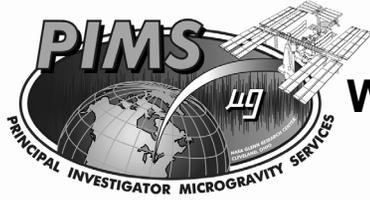




EXPERIMENT SENSITIVITY ASSESSMENT

Measurement Needs By Disciplines

Experiment Type	Frequency Range	Measurement Level
Biotechnology	QS – 10 Hz	100 µg and above
Fluid Physics	QS – 300 Hz	1 µg to 1 mg
Combustion Science	QS – 50 Hz	10 µg and above
Fundamental Physics	QS – 180 Hz	0.1 µg and above
Material Science	QS – 300 Hz	0.01 µg and above



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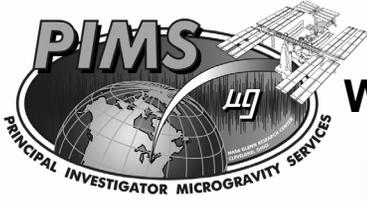
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--- Overview**
- ▶ **ISS Microgravity Environment Requirements**
- ▶ **Microgravity Disciplines Sensitivity Assessment**
- ▶ **Vibration Isolation--- Why?**
- ▶ **ISS Coordinates Systems**
- ▶ **Plots Header Description: ISS**
- ▶ **Experiment Planning and Execution**
- ▶ **Preflight Planning for Science Optimization**
- ▶ **Overall Summary**
- ▶ **Extra Information: STS Ascent and Landing Profiles**



Vibration Isolation

MOTIVATION

- Ambient spacecraft acceleration levels often are higher than allowable from a science perspective, therefore:
 - Vibration isolation is used to reduce the acceleration levels to an acceptable level
 - Either passive or active isolation is used depending on the needs or requirements of a specific facility or experiments
 - These vibration isolation systems have flown on the Space Shuttle:
 - STABLE, MIM, and ARIS RME
 - The following vibration isolation systems are (or will be) on the ISS:
 - ARIS, MIM-2, g-LIMIT



Vibration Isolation Systems

Vibration Isolation

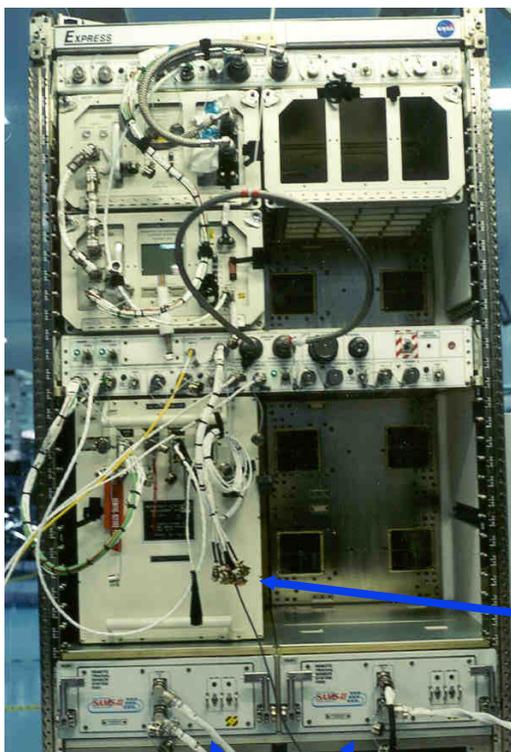
- **Rack Level Isolation Systems:**
 - **Active Rack Isolation System– ARIS**
 - ARIS provides rack-level reduction of acceleration levels
 - ARIS supplied by ISSP to meet the microgravity requirements
 - 6 racks will have ARIS installed
 - **Passive Rack Isolation System-- PaRIS**
 - ◇ Provides passive isolation from 0.5 - 2.0 Hz. Up to 300 Hz.
 - ◇ Will be used on 3 racks: 2 HHRs and CIR
- **Sub-rack Isolation Systems**
 - **STABLE / g LIMIT (Glovebox)**
 - Marshall Space Flight Center
 - **MIM/ MIM-2**
 - Canadian Space Agency

ISS Rack Facilities

EXPRESS RACK #1

Increment-2 Configuration

EXPRESS RACK #2



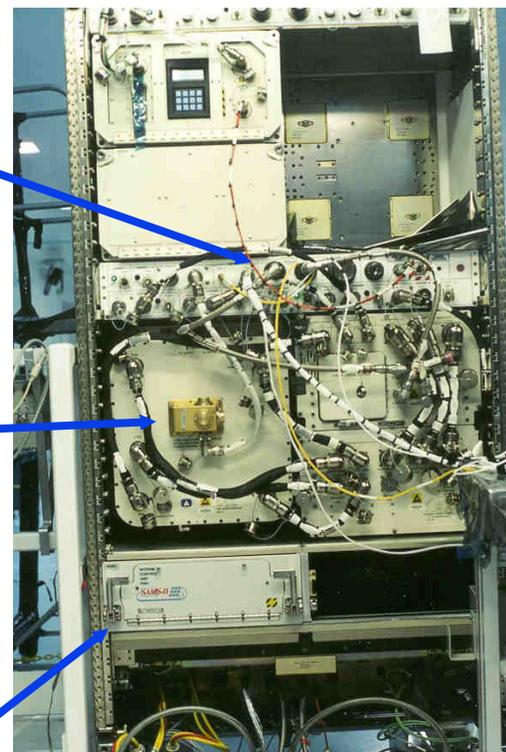
RTS DRAWERS

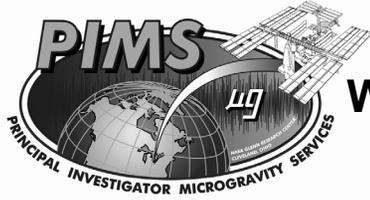
RTS-EE installed in EXPRESS rack behind panel (standard in ARIS EXPRESS racks)

RTS-SE

MAMS

ICU DRAWER



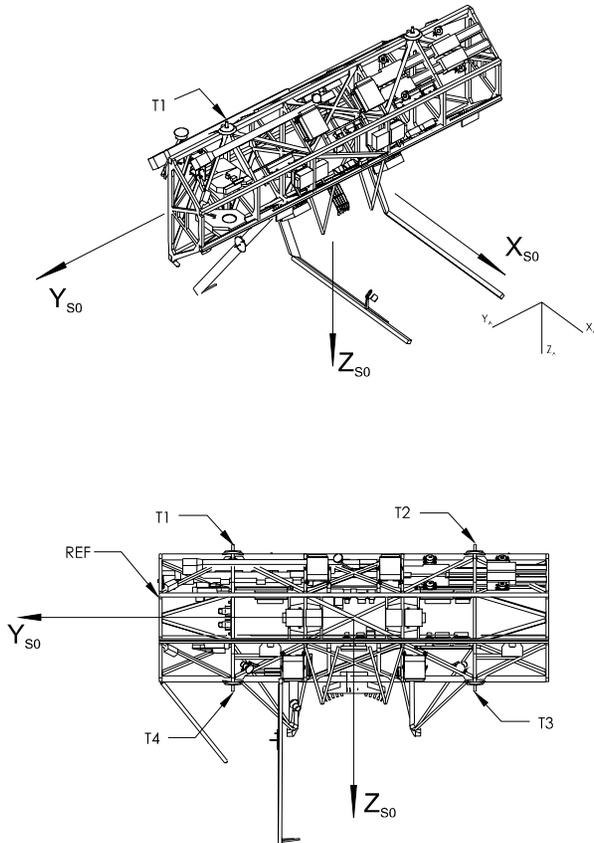


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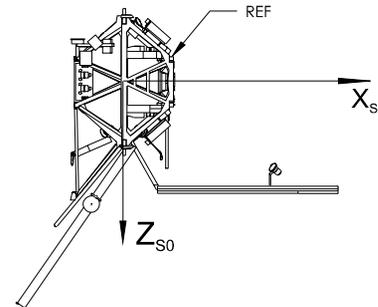
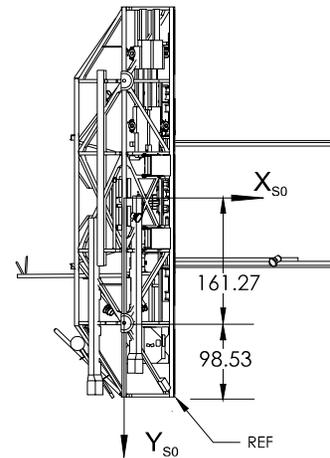
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--- Overview**
- ▶ **ISS Microgravity Environment Requirements**
- ▶ **Microgravity Disciplines Sensitivity Assessment**
- ▶ **Vibration Isolation--- Why?**
- ▶ **ISS Coordinates Systems**
- ▶ **Plots Header Description: ISS**
- ▶ **Experiment Planning and Execution**
- ▶ **Preflight Planning for Science Optimization**
- ▶ **Overall Summary**
- ▶ **Extra Information: STS Ascent and Landing Profiles**

COORDINATE SYSTEMS

ISS



Integrated Truss Segment S0 Coordinate System



Type

Right-Handed Cartesian, Body-Fixed

Description

This coordinate system defines the origin, orientation, and sense of the Space Station Analysis Coordinate System.

Origin

The YZ plane nominally contains the centerline of all four trunnion pins. The origin is defined as the intersection of two diagonal lines connecting the centers of the bases of opposite trunnion pins, running T1 to T3 and from T2 to T4.

Orientation

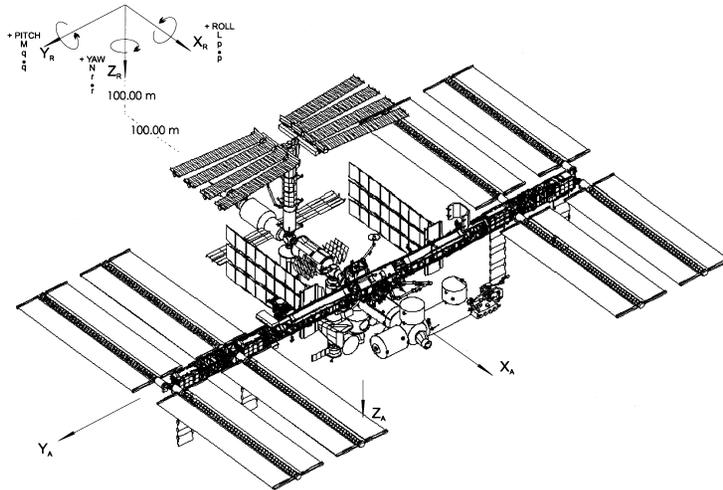
X_{S0}: The X-axis is parallel to the vector cross-product of the Y-axis with the line from the center of the base trunnion pin T2 to the center of the base trunnion pin T3, and is positive forward

Y_{S0}: The Y-axis is parallel with the line from the center of the base of trunnion pin T2 to the center of the base of trunnion pin T1. The positive Y-axis is toward starboard.

Z_{S0}: The Z-axis completes the RHCS

COORDINATE SYSTEMS

ISS



SPACE STATION ANALYSIS COORDINATE SYSTEM

Type

Right-Handed Cartesian, Body-Fixed

Description

This coordinate system is derived using the Local Vertical Local Horizontal (LVLH) flight orientation. When defining the relationship between this coordinate system and another, the Euler angle sequence to be used is a yaw, pitch, roll sequence around the Z_A, Y_A, and X_A axes, respectively.

Origin

The origin is located at the geometric center of Integrated Truss Segment (ITS) S₀ and is coincident with the S₀ Coordinate frame.

Orientation

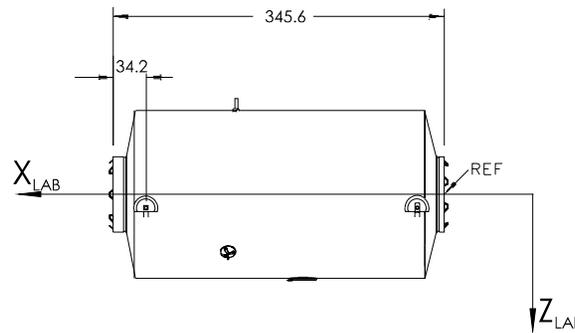
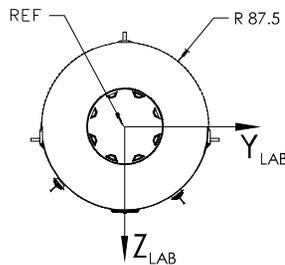
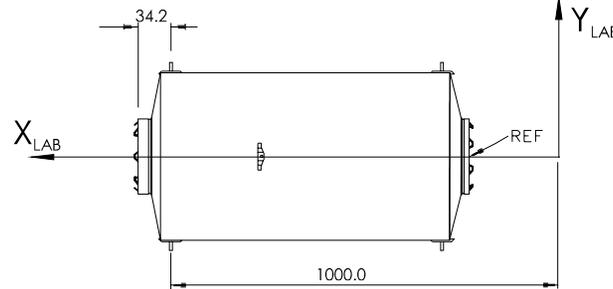
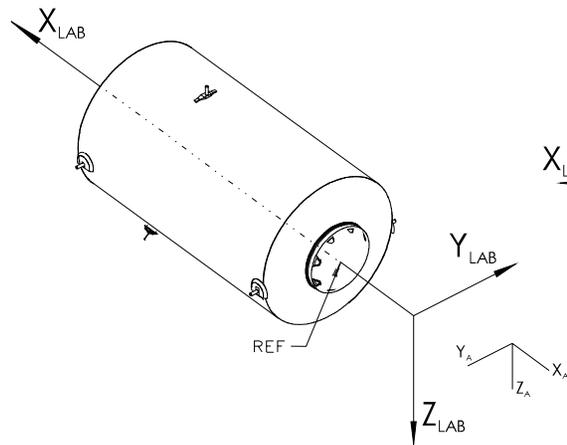
X_A: The X-axis is parallel to the longitudinal axis of the module cluster. The positive X-axis is in the the forward direction

Y_A: The Y-axis is identical with the S₀ axis. The nominal alpha joint rotational axis is parallel with Y_A. The positive Y-axis is in the starboard direction.

Z_A: The positive Z-axis is in the direction of nadir and completes the right-handed Cartesian system (RHCS).

COORDINATE SYSTEMS

ISS



UNITED STATES LABORATORY MODULE COORDINATE SYSTEM

Type

Right-Handed Cartesian, Body-Fixed to the Pressurized Module

Origin

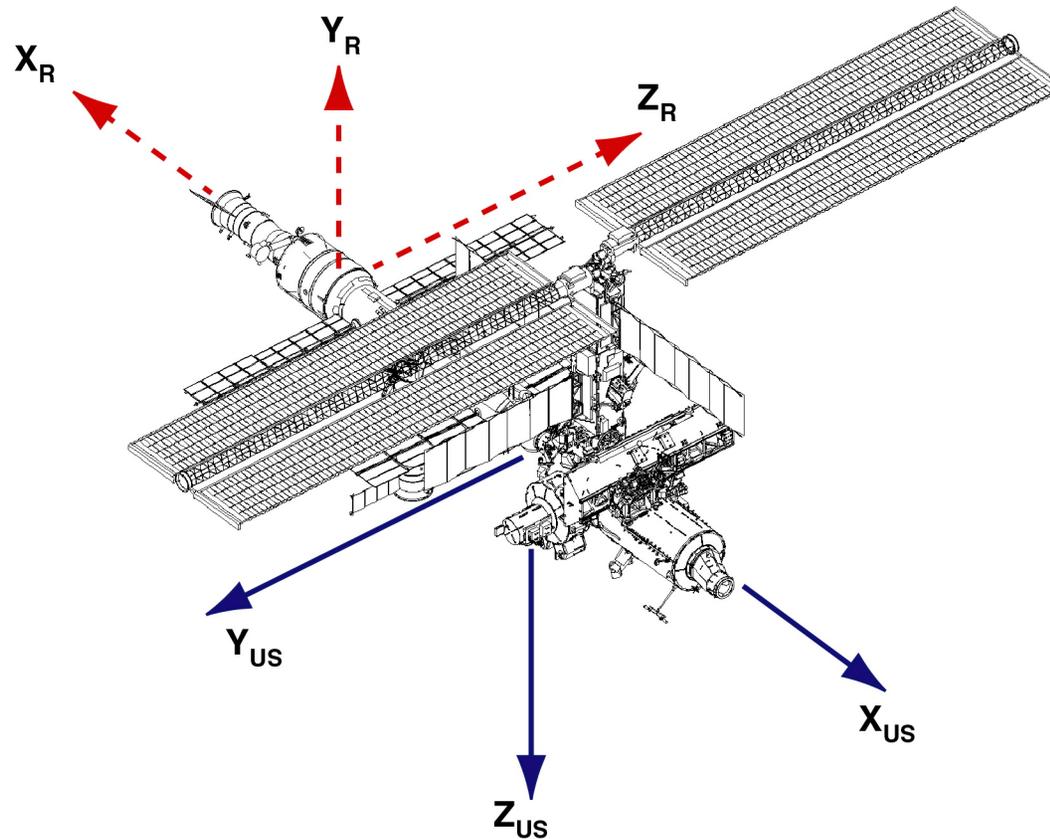
The origin is located forward of the pressurized module such that the center of the bases of the aft trunnions have X_{LAB} components nominally equal to 1000.000 inches.

Orientation

X_{LAB} : The X-axis is perpendicular to the nominal aft CBM interface plane and pierces the geometric center of the array of mating bolts at the aft end of the pressurized module. The positive X-axis is toward the pressurized module from the origin.

Y_{LAB} : The Y-axis completes the right-handed Cartesian system (RHCS).

Z_{LAB} : The Z-axis is parallel to the perpendicular line from the X-axis to the center of base of the keel pin, and positive in the opposite direction as shown.



ISS-USOS / ROS Analysis Coordinate Systems

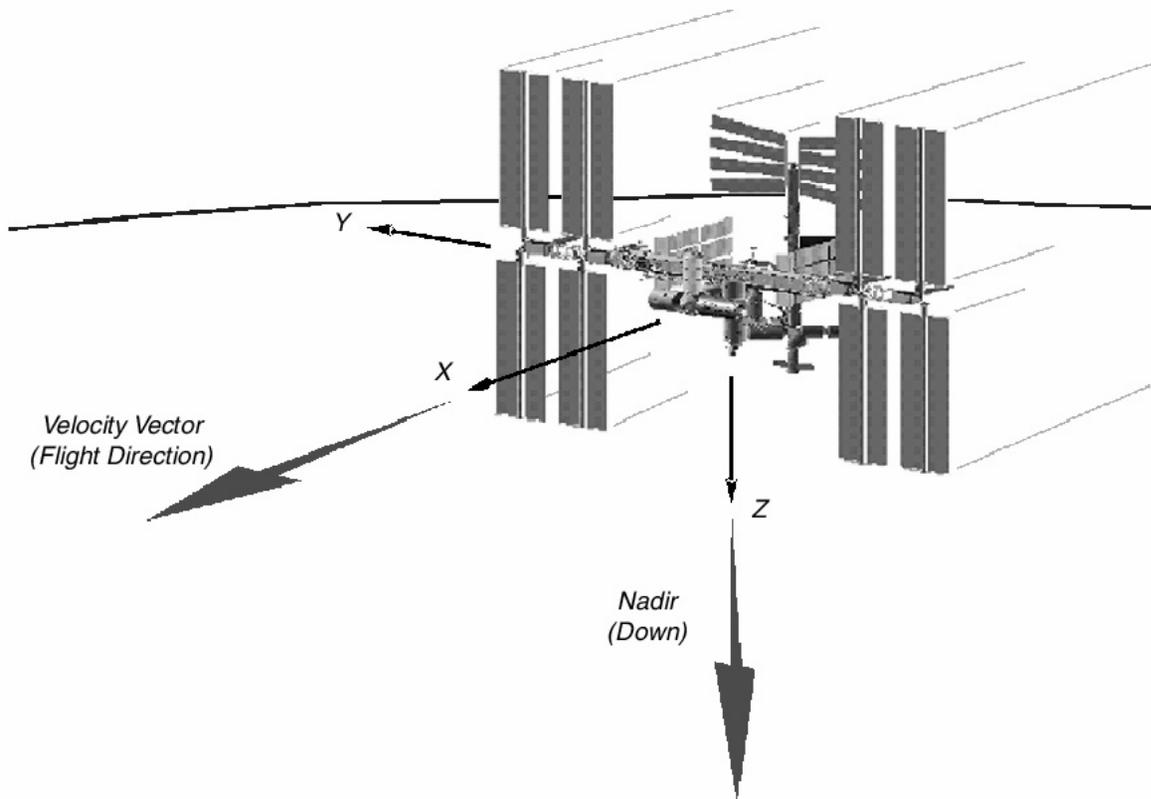
FLIGHT ATTITUDES

ISS-XVV Z Nadir

XVV Z Nadir: X Axis Near Velocity Vector, Z Axis Nadir/Down

XVV Z Nadir Flight Attitude Shown With 0, 0, 0 Deg. Yaw, Pitch, Roll LVLH Attitude

XVV TEA is Nearest Torque Equilibrium Attitude (TEA) To This Orientation



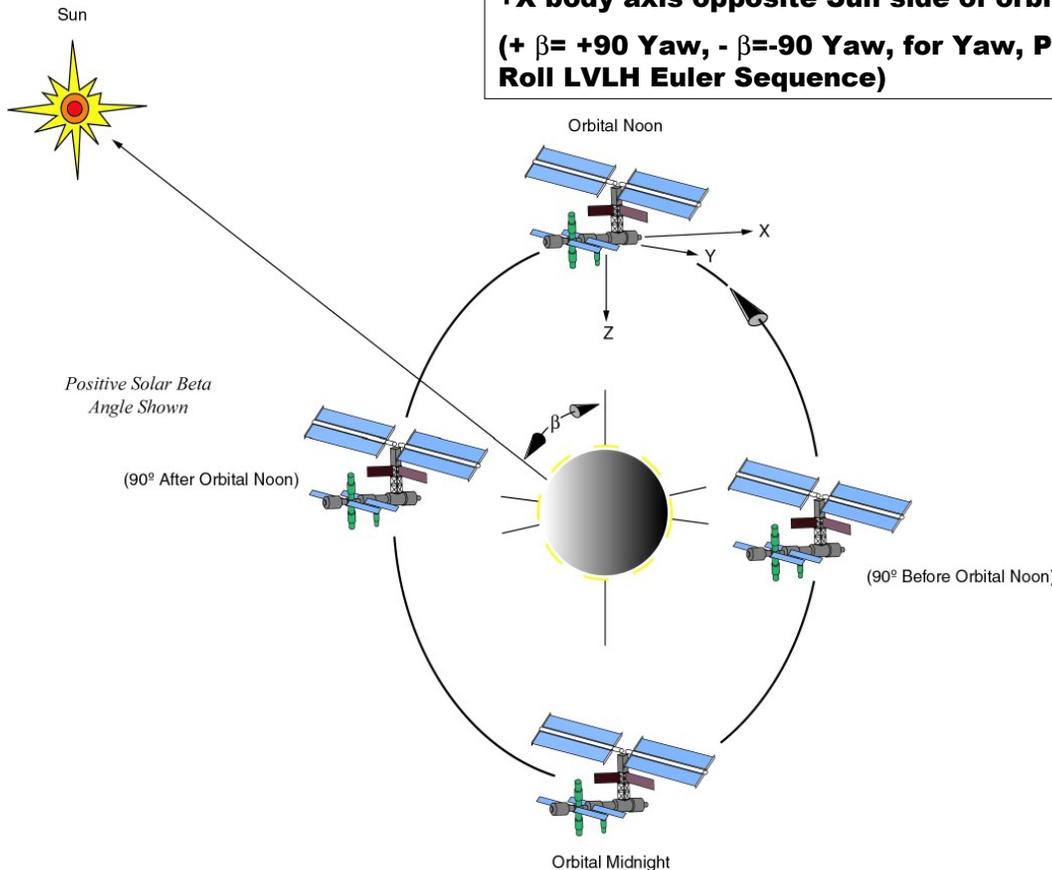
The basic flight attitude for ISS is called **XVV Z Nadir**. The vehicle design is optimized for this attitude. The **XVV** attitude:

- places the most modules in the microgravity volume
- supports altitude reboosts
- service vehicle dockings
- minimizes aerodynamic drag

The ISS is designed to tolerate deviations from perfect XVV Z Nadir of +/- 15 degrees in each axis. This envelope was expanded to -20 deg in pitch.

FLIGHT ATTITUDES

ISS-XPOP

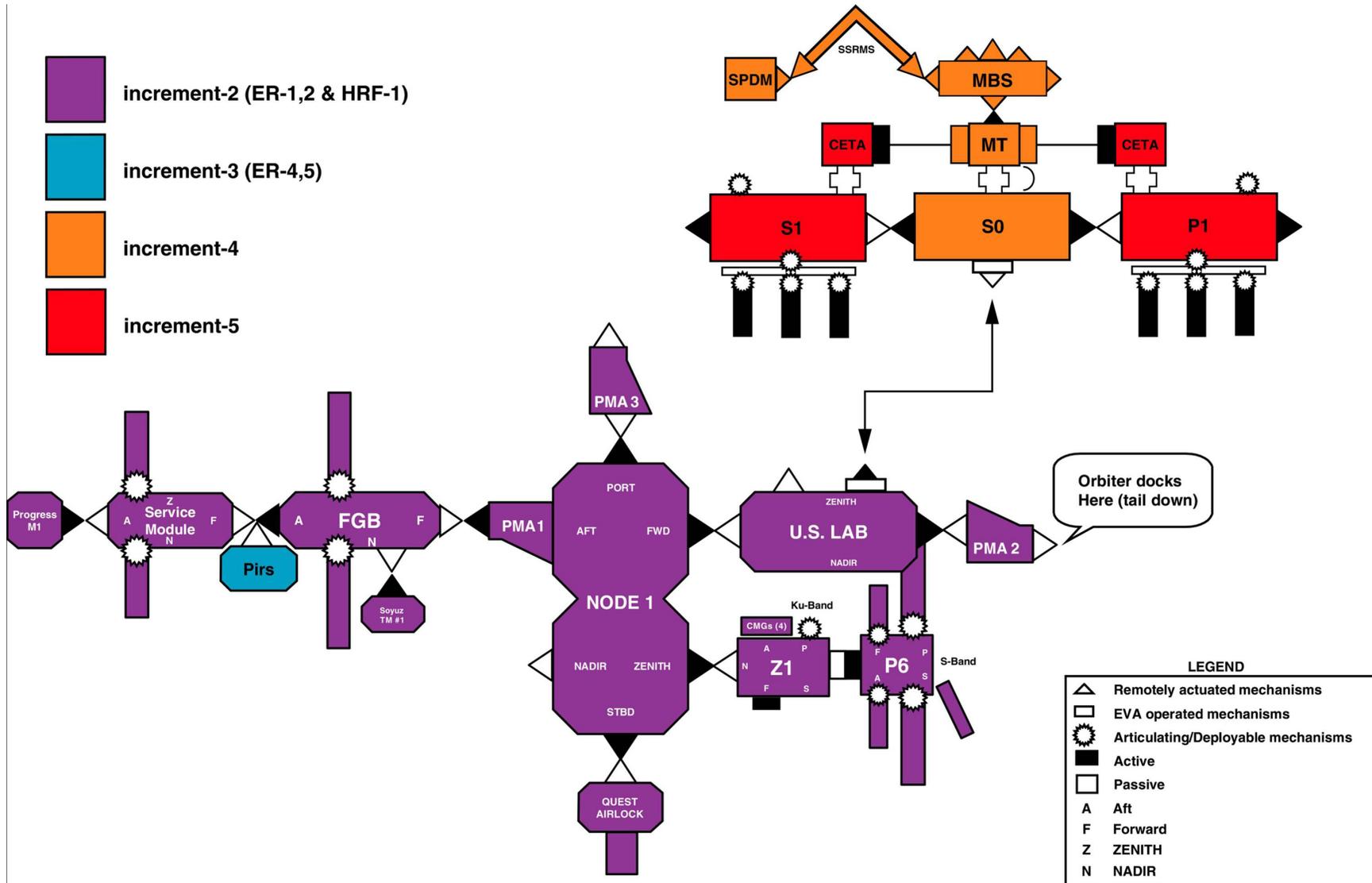


+ Z body axis is down/Nadir at orbital noon
+X body axis opposite Sun side of orbit plane
(+ $\beta = +90$ Yaw, - $\beta = -90$ Yaw, for Yaw, Pitch, Roll LVLH Euler Sequence)

During the assembly stages (stages 2A through 12A.1), **ISS** will not be capable of generating enough power to sustain the required electrical loads in the **XVV** flight attitude at mid-to-high solar beta angles because these vehicle configurations have only a single solar array gimbal axis, which is aligned so that it only perfectly tracks the Sun when the solar beta angle is near zero. Therefore, **ISS** is designed to accommodate a second basic flight orientation for these increments.

XPOP: X-principal axis perpendicular to the orbit plane, Z Nadir at orbital noon. The **XPOP** flight attitude sets up geometry between the ISS and the Sun so that the Sun stays close to the ISS/XZ body axis plane. This allows all the solar arrays to track the Sun regardless of the solar beta angle. **XPOP** also places the dominant inertia axis in the local horizontal to minimize gravity gradient torques and allow Control Moment Gyro (CMG) non-propulsive attitude control.

- **Maximum power generation**
- **Minimizes vehicle gravity gradient torque**
- **Momentum manager provides CMG attitude control without propellant usage**





CONTENT

- ▶ **Reduced Gravity Environment Description
--- Overview**
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- ▶ **Vibration Isolation--- Why?**
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- ▶ **Plots Header Description: ISS**
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- ▶ **Overall Summary**
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SOME DEFINITIONS

Acceleration Measurement Systems

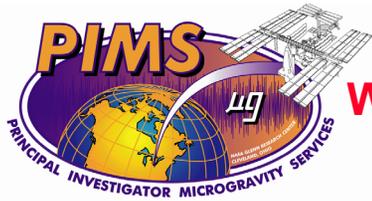
- **MAMS-OSS:** Microgravity Acceleration Measurement System – OARE SubSystems- instrument which measures the quasi-steady acceleration levels to characterize the ISS reduced gravity environment provided to users. MAMS measures accelerations from DC to 1 Hz. However, PIMS analyzes and reports the data up to 0.01 Hz.
- **MAMS-HiRAP:** Microgravity Acceleration Measurement System- High Resolution Acceleration Package- instrument which measures the vibratory accelerations from 0.01 to 100 Hz aboard the ISS
- **SAMS (RTS):** Second generation SAMS - instrument which measures the vibratory accelerations level from 0.01 to 300 Hz aboard the ISS.
- **SAMS (TSH):** - instrument for stand alone missions (e.g. sounding rockets), Shuttle, drop towers and KC-135, which measures vibratory acceleration levels from 0.01 to 200 Hz.



SOME DEFINITIONS

Definitions

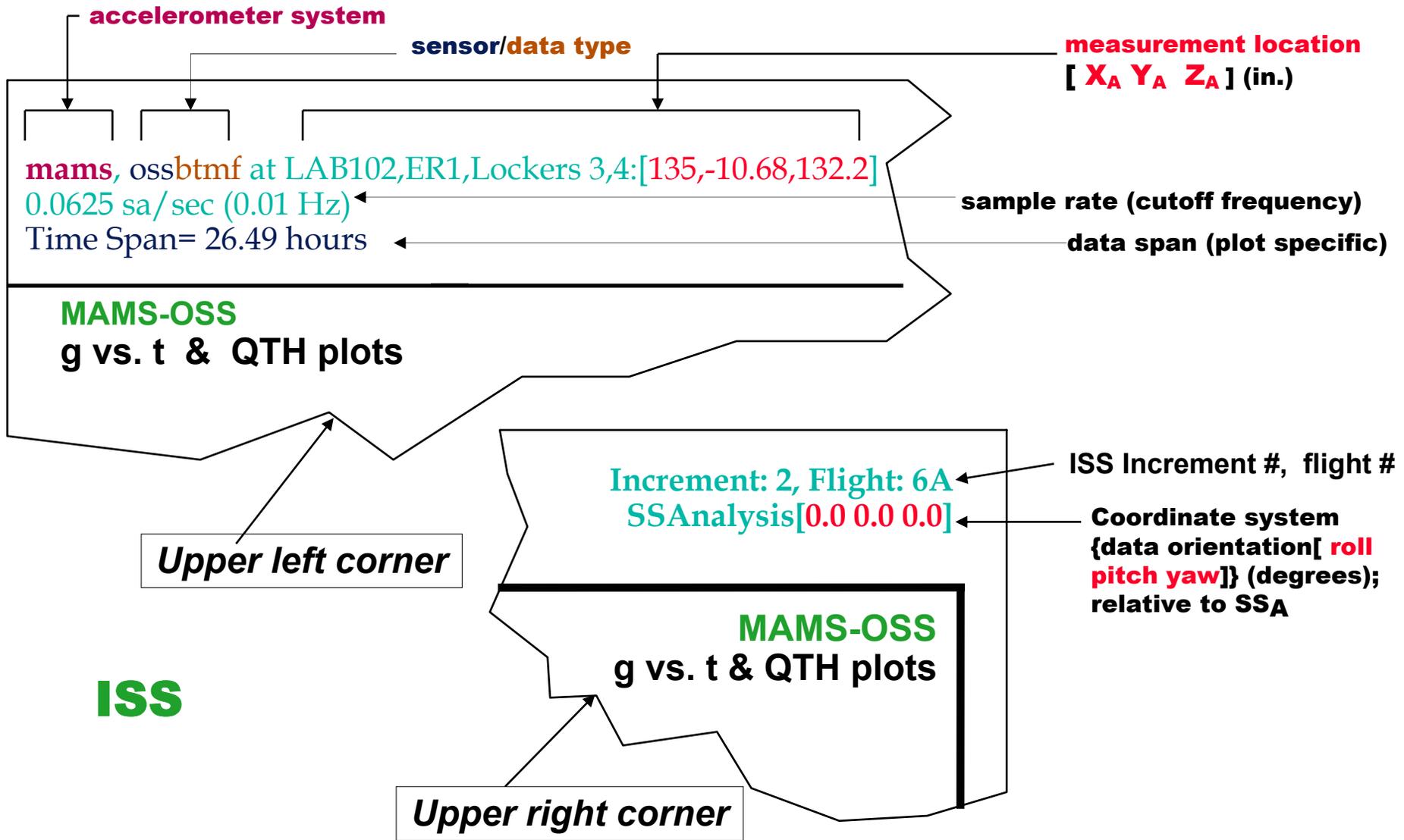
- **Nyquist criteria:** sampling rate must be at least twice that of the highest frequency contained in the signal of interest
- **Cutoff frequency (f_c):** corner frequency in filter response; highest unfiltered frequency of interest
- **Sample rate (f_s):** rate at which an analog signal is sampled ($\text{samples}/\text{sec}$)
- **Power spectral density:** a function that quantifies the distribution of power in a signal with respect to frequency
- **Spectrogram:** a 3-D representation of the power spectral density as a function of frequency and time

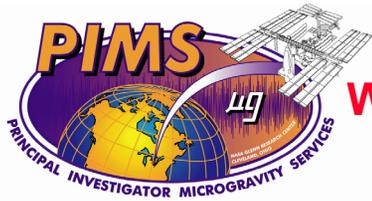


Working in a Reduced Gravity Environment : "A Primer"



ISS QUASI-STEADY SAMPLE PLOTS INFORMATION

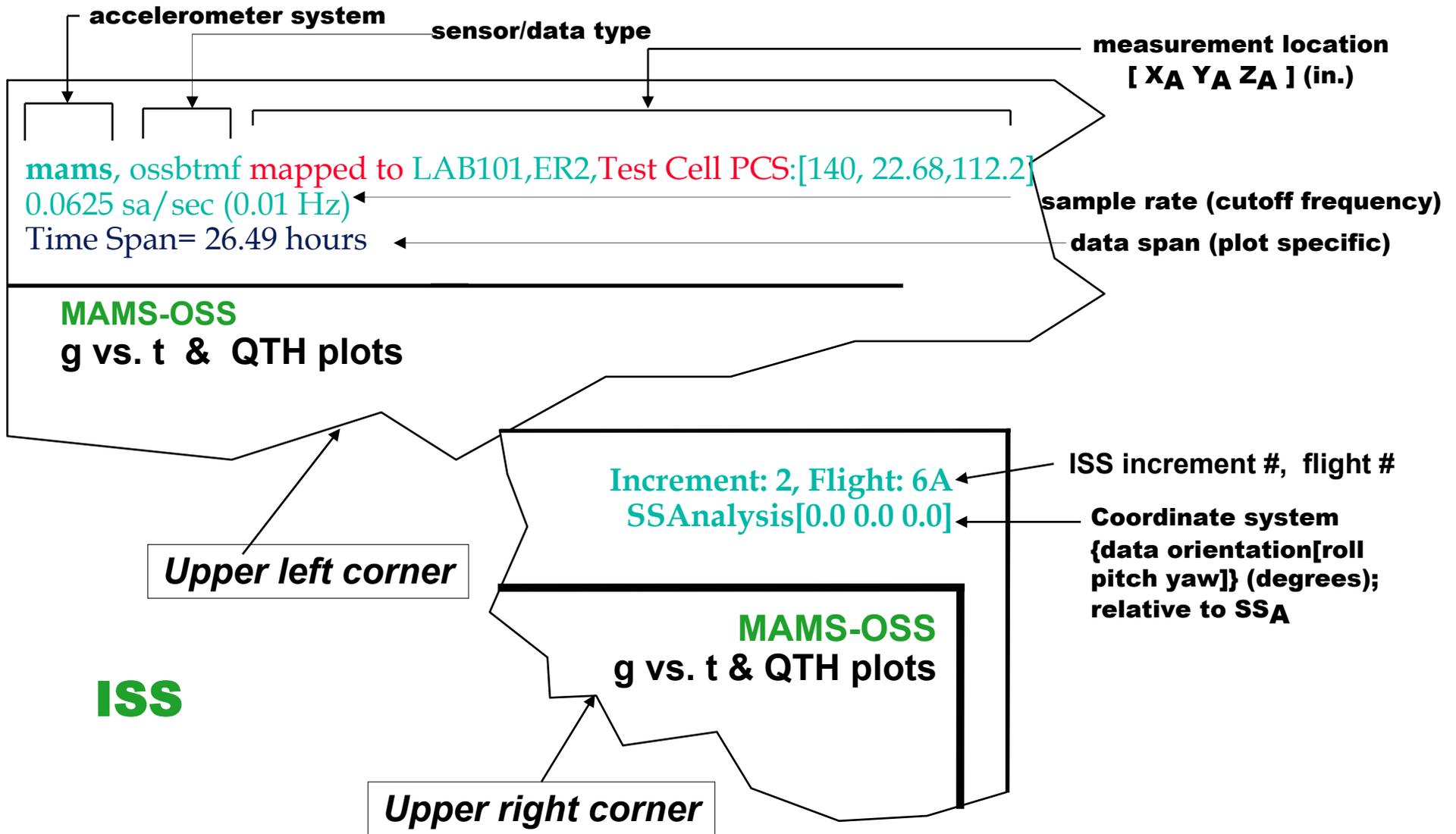


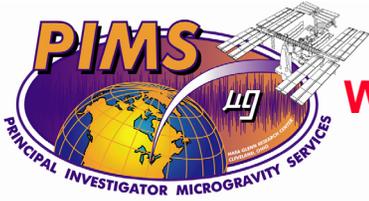


Working in a Reduced Gravity Environment : "A Primer"



ISS QUASI-STEADY SAMPLE PLOTS INFORMATION

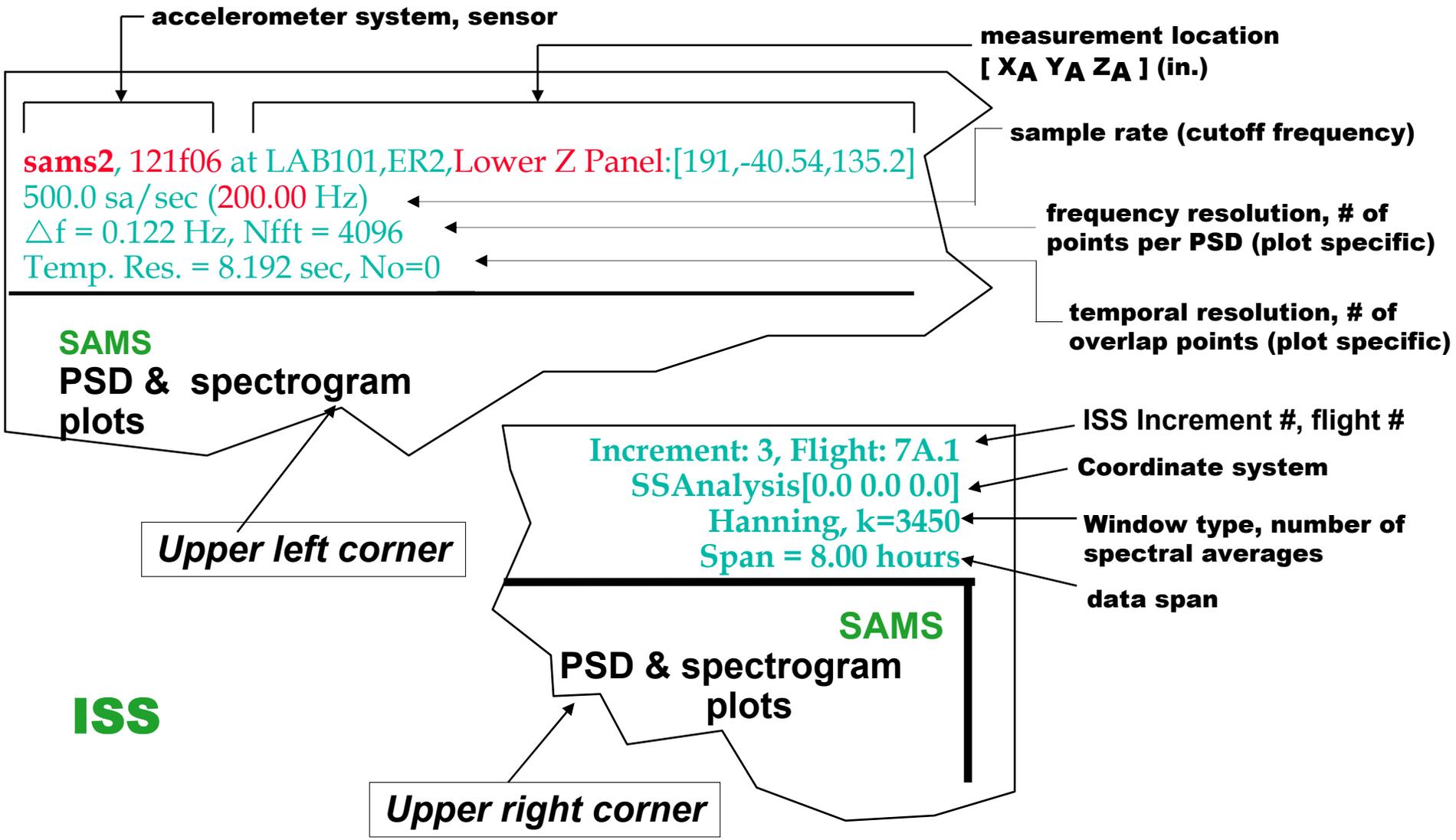


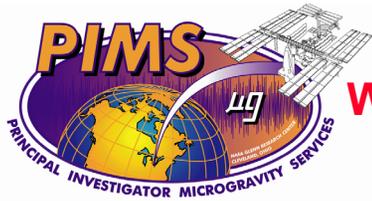


Working in a Reduced Gravity Environment : "A Primer"



ISS VIBRATORY SAMPLE PLOTS INFORMATION





Working in a Reduced Gravity Environment : “A Primer”



sams2, 121f03 at LAB1O1, ER2, Lower Z Panel:[191.54 -40.54 135.25]

1000.0 sa/sec (400.00 Hz)

$\Delta f = 0.488$ Hz, Nfft = 2048

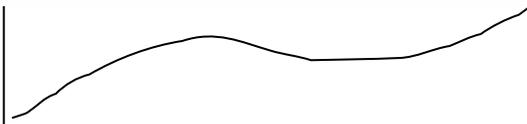
Temp. Res. = 2.048 sec, No = 0

Increment: 4, Flight: UF1

SSAnalysis[0.0 0.0 0.0]

Hanning, k = 586

Span = 19.97 minutes

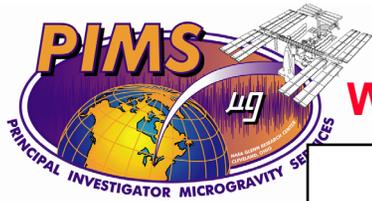


Upper Left Corner:

- Line 1
 - accelerometer system, sensor
 - measurement location [x_A y_A z_A] (in.)
- Line 2
 - sample rate (samples per second)
 - cutoff frequency (Hz)
- Lines ≥ 3 are plot specific
 - frequency resolution
 - temporal resolution

Upper Right Corner:

- Line 1
 - ISS Increment & Flight
- Line 2
 - data orientation [roll pitch yaw] (degrees); relative to SSA
- Lines ≥ 3 are plot specific
 - window type
 - number of spectral averages
 - data span



Working in a Reduced Gravity Environment : "A Primer"



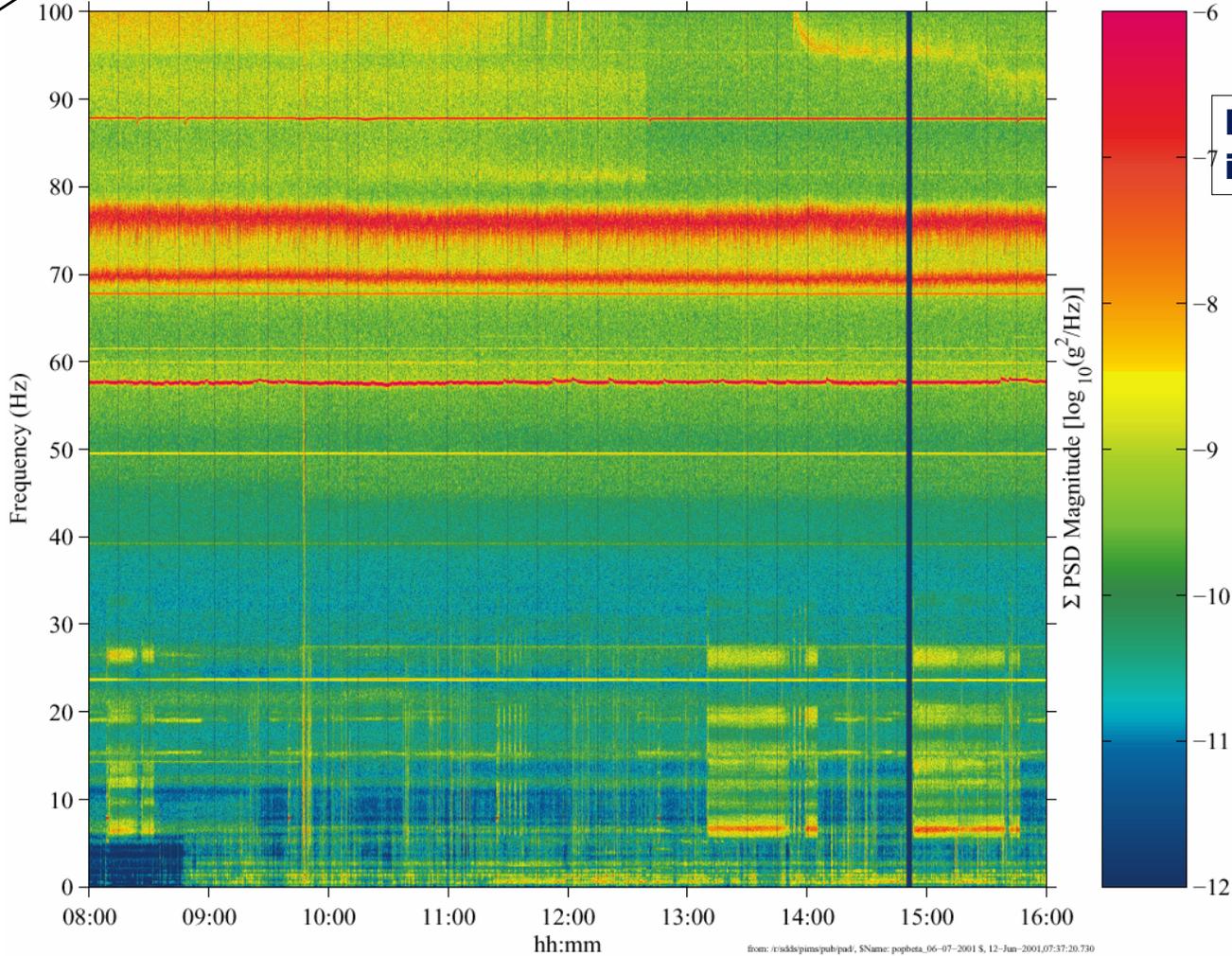
Plot Header Sample

mams, hirap at LAB1Q2, ER1, Lockers 3,4:[138.68 -16.18 142.35]
 1000.0 sa/sec (100.00 Hz)
 $\Delta f = 0.122$ Hz, Nfft = 8192
 Temp. Res. = 8.192 sec, No = 0

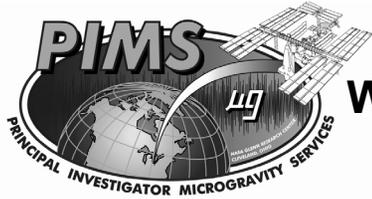
MAMS HiRAP
 Start GMT 2001:06:03:08:00:00

Increment: 2, Flight: 6A
 Sum
 Hanning, k = 3,6
 Span = 8.00 hours

Plot header information

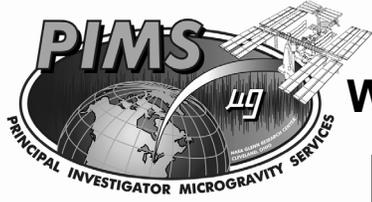


Plot header information



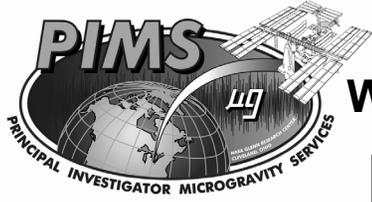
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- ▶ **Reduced Gravity Environment Description
--- Overview**
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- ▶ **Microgravity Disciplines Sensitivity Assessment**
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Available Reduced Gravity Carriers / Facilities

- **STS Orbiters**
- **International Space Station (ISS)**
- **Sounding Rockets (Various Countries)**
- **Parabolic Flight Aircraft (KC-135)**
- **Free-Flyers (ESA-Russia)**
- **Drop Towers (US, Germany, Japan)**
- **Microgravity Emission Lab (MEL) @ NASA-GRC**



Experiment Location and Orientation

- **Proximity to carrier / vehicle center of mass**
 - sensitivity to quasi-steady variations
- **Proximity to other equipment**
 - sensitivity to vibration sources
- **Alignment**
 - sensitivity to quasi-steady acceleration direction



EXPERIMENT PLANNING AND EXECUTION

Carrier Attitude

- **Issues related to experiment location**
 - gravity gradient effects
- **Issues related to experiment orientation**
 - design attitude that points experiment in desired direction
- **Sensitivity to quasi-steady variations with time**
 - atmospheric drag effects
 - local vertical / local horizontal attitudes versus inertial attitude



Accelerometer Selection

- **Frequency Range**
 - **cutoff frequency based on experiment sensitivity**
 - **sampling rate and filter characteristics specified by accelerometer system team to provide frequency selected by experimenter**
- **Location and Alignment**
 - **close to experiment sensitive location**
 - **mounting technique**
 - **away from sources which may disturb accelerometer and mask disturbances of interest**
 - **knowledge of sensor orientation relative to experiment axes**

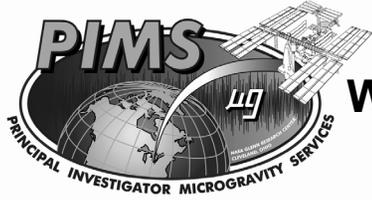


EXPERIMENT PLANNING AND EXECUTION

Mission / Experiment Timeline

If at all possible, schedule your experiment operations to avoid any activities which might negatively impact it. Keep the following points in mind:

- **Experiment sensitivity to acceleration sources**
 - quasi-steady, vibratory and transient
- **Crew exercise**
- **Crew activity**
- **Thruster activity**
- **Other experiment operations**
- **Venting**



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Motivation

- **Real-time replanning requires clear goals to be effective**
- **Decision makers may have no time to deeply consider the impacts in a crisis situation (it is after all your job– it is your experiment)**
- **PIs will fare better if they have a cohesive and rational plan (for-off nominal conditions)**



Potential Success-level Definition

- **Technology demonstration:**

- The level of success necessary to validate the hardware functionality and to observe critical science concepts (e.g. approach to steady state). The concept is that any follow-on experiment would benefit substantially from this level of testing.

- **Minimum Science:**

- The level of success necessary to produce a scientifically interesting paper that is publishable in an important journal (or to meet minimal commercial goals).

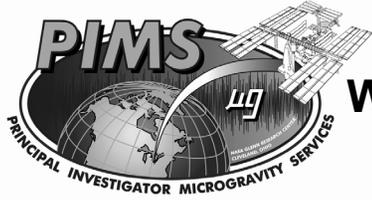
- **Minimal achievement of experiment objectives**

- The level of success necessary to satisfy a minimal number of the peer reviewed experiment objectives (or approved commercial goals). This is typically the level which, if it can not be assured, the hardware developer would not ship the hardware.



Planning Goals

- **Plan for at least minimum science (in extreme situations)**
- **Plan to reach minimal achievement of the experiment objectives (in less extreme situations)**
- **Allow for equitable distribution of the timeline for both reduced and extra cases**
- **Write clear, concise, logical procedures (step by step) for the crew (astronauts) to follow, if crew is required**



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OVERALL SUMMARY

- **The reduced gravity environment is not “zero-g” or even “zero-acceleration”. It is dynamic.**
- **The carrier environment may (and will) influence the results of a science experiment:**
 - **Carrier hardware**
 - **Experiment hardware**
 - **Crew effects**
 - **Water dump / Venting**
 - **Carrier attitude**
 - **Carrier altitude**
 - **Jet firings**



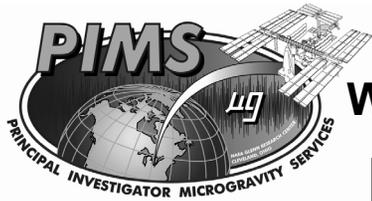
OVERALL SUMMARY

- **Analyses and/or tests should be performed before flight to investigate the sensitivity of an experiment to the reduced gravity environment.**
- **Environments of past missions should be considered in planning future experiments (PIMS is a good source for that)**
- **Experiment teams MUST understand their own experiment hardware both for sensitivities and potential disturbance sources they may be causing to the environment with (for example) moving parts from their experiments or / and required crew actions (observe the “good neighbor policy”)**



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Working in a Reduced Gravity Environment : "A Primer"

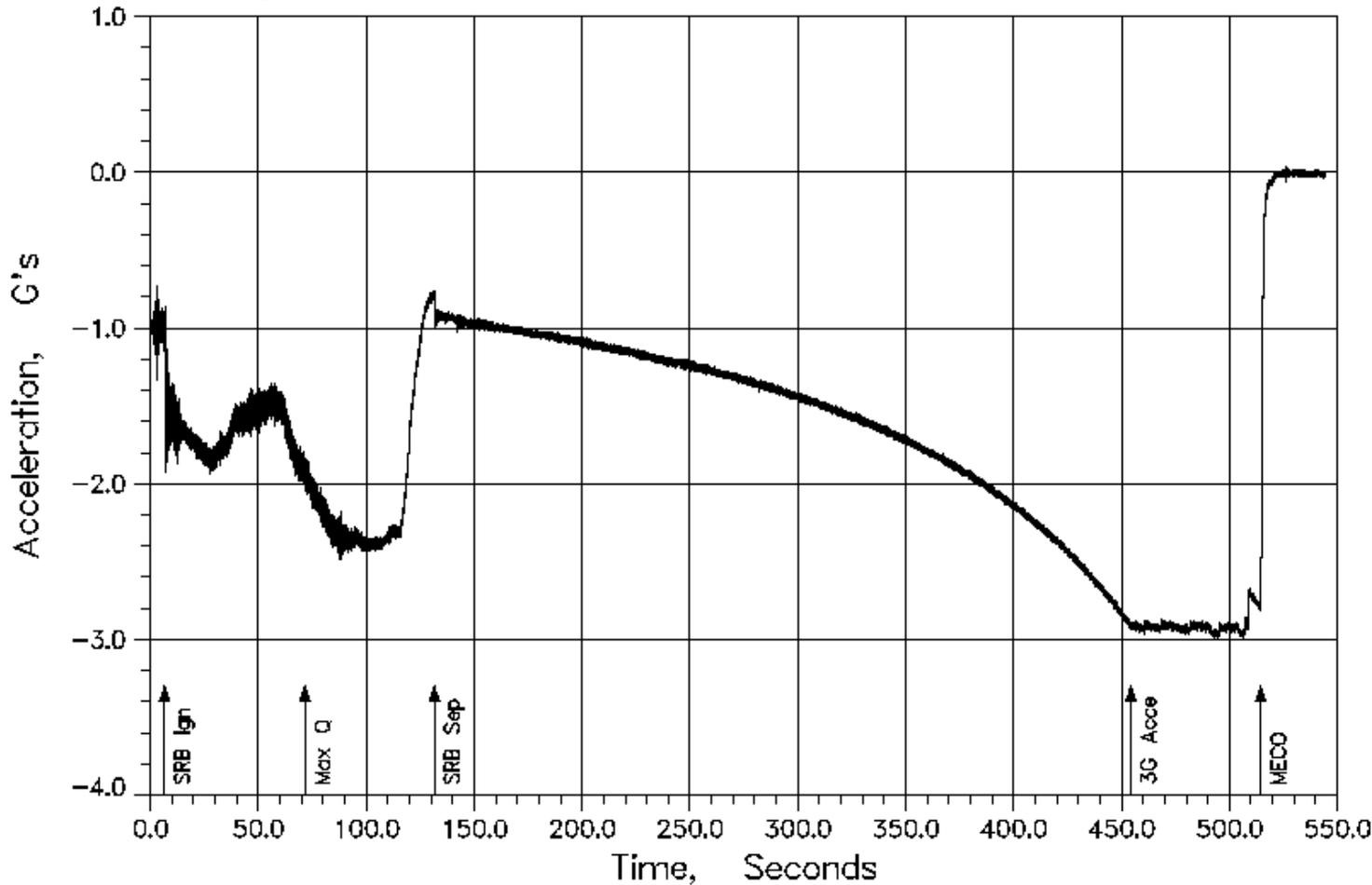


STS ASCENT PROFILE

STS-90 Payload Bay Ascent Data - SSME Ign=0.0 second (0-50 Hz)

V34A9483A, DOF: X, Location: x=1029.0, y=-101.0, z=408.0

SRB-Ign=6.56, Max-Q=71.57, SRB-Sep=131.57, 3G-Acce=454.37, MECO=514.57



STS = Columbia

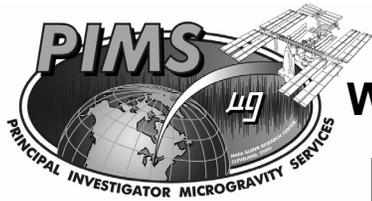
SRB_{ign} = solid rocket boosters ignition

Max Q = time of maximum dynamic pressure

SRB_{Sep} = solid rocket boosters separation

3G Acce = time at which 3g acceleration is reached

MECO = main engine cutoff



Working in a Reduced Gravity Environment : "A Primer"

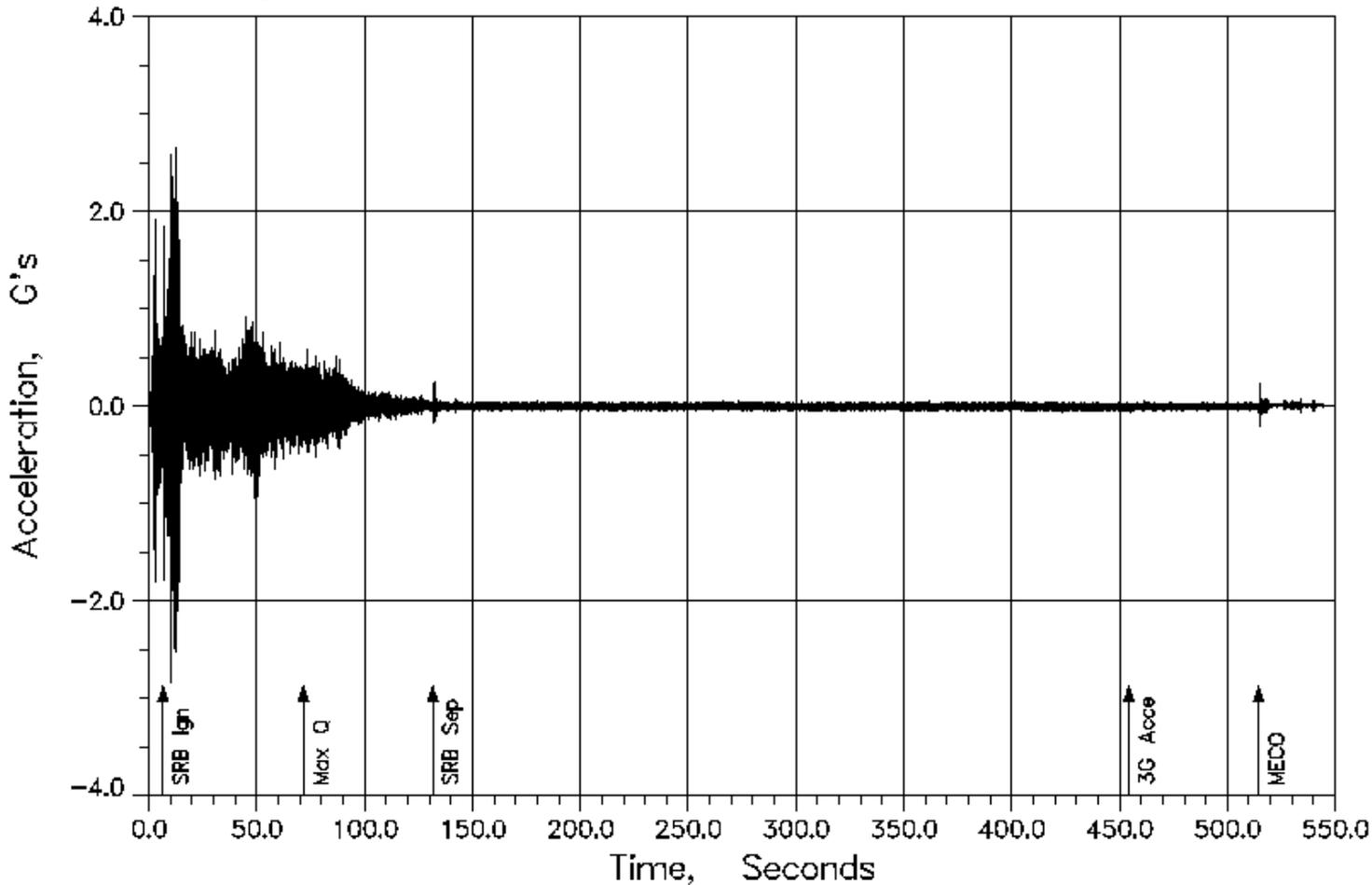


STS ASCENT PROFILE

STS-90 Payload Bay Ascent Data - SSME Ign=0.0 second (0-50 Hz)

V34A9460A, DOF: Y, Location: x=701.0, y=-102.0, z=407.0

SRB-Ign=6.56, Max-Q=71.57, SRB-Sep=131.57, 3G-Acce=454.37, MECO=514.57



- **Sensor dynamic responses:**
-10 to + 10 G range
- **Frequency response:**
0 – 50 Hz
- **Sampling rate:**
500 samples per sec.
- **Coordinate system:**
Orbiter structural system



Working in a Reduced Gravity Environment : "A Primer"

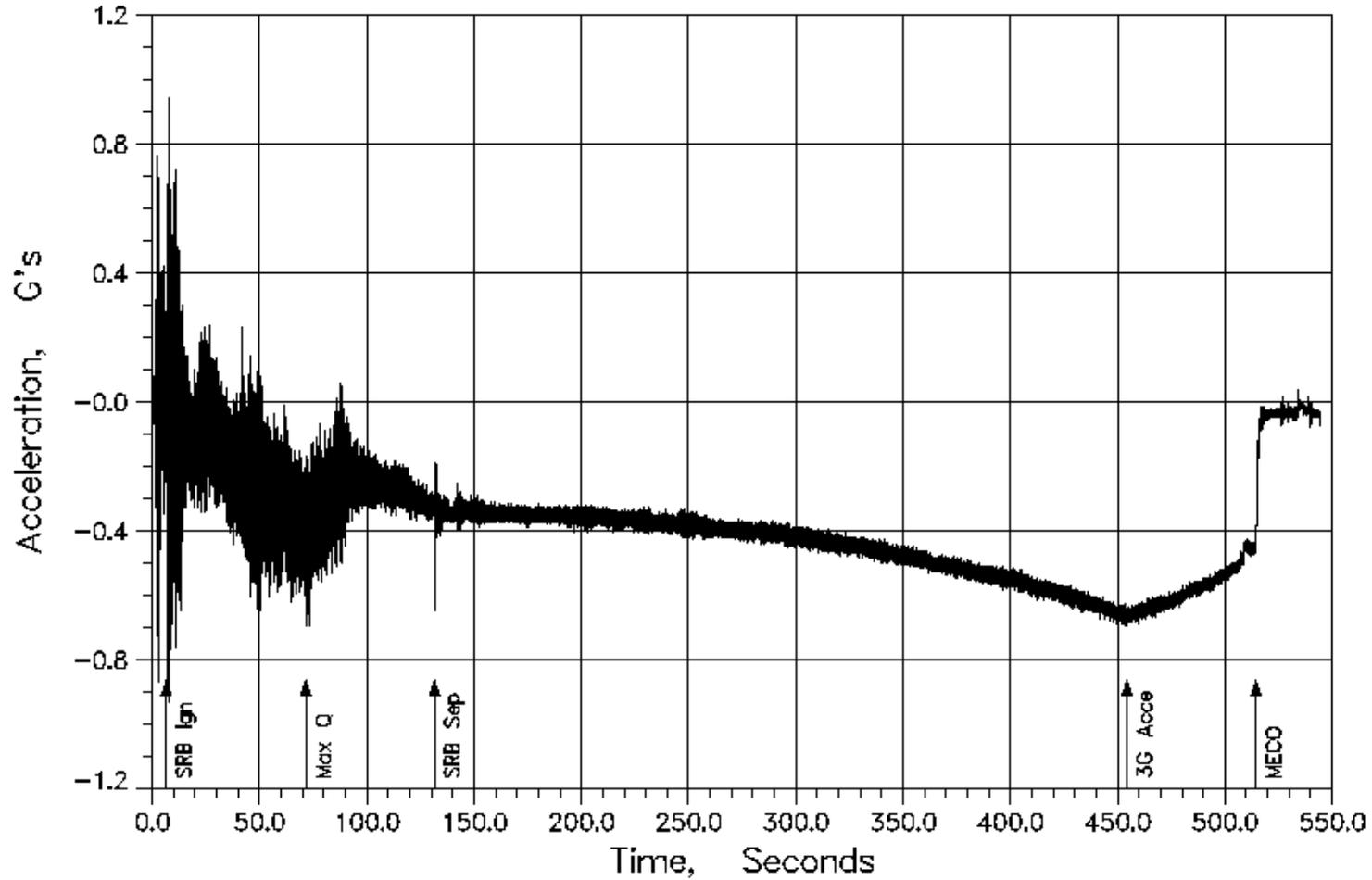


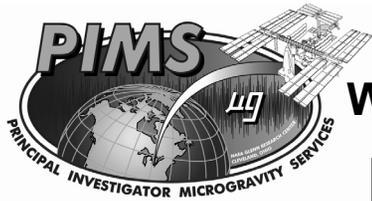
STS ASCENT PROFILE

STS-90 Payload Bay Ascent Data - SSME Ign=0.0 second (0-50 Hz)

V34A9461A, DOF: Z, Location: x=701.0, y=-102.0, z=407.0

SRB-Ign=6.56, Max-Q=71.57, SRB-Sep=131.57, 3G-Acce=454.37, MECO=514.57





Working in a Reduced Gravity Environment : "A Primer"

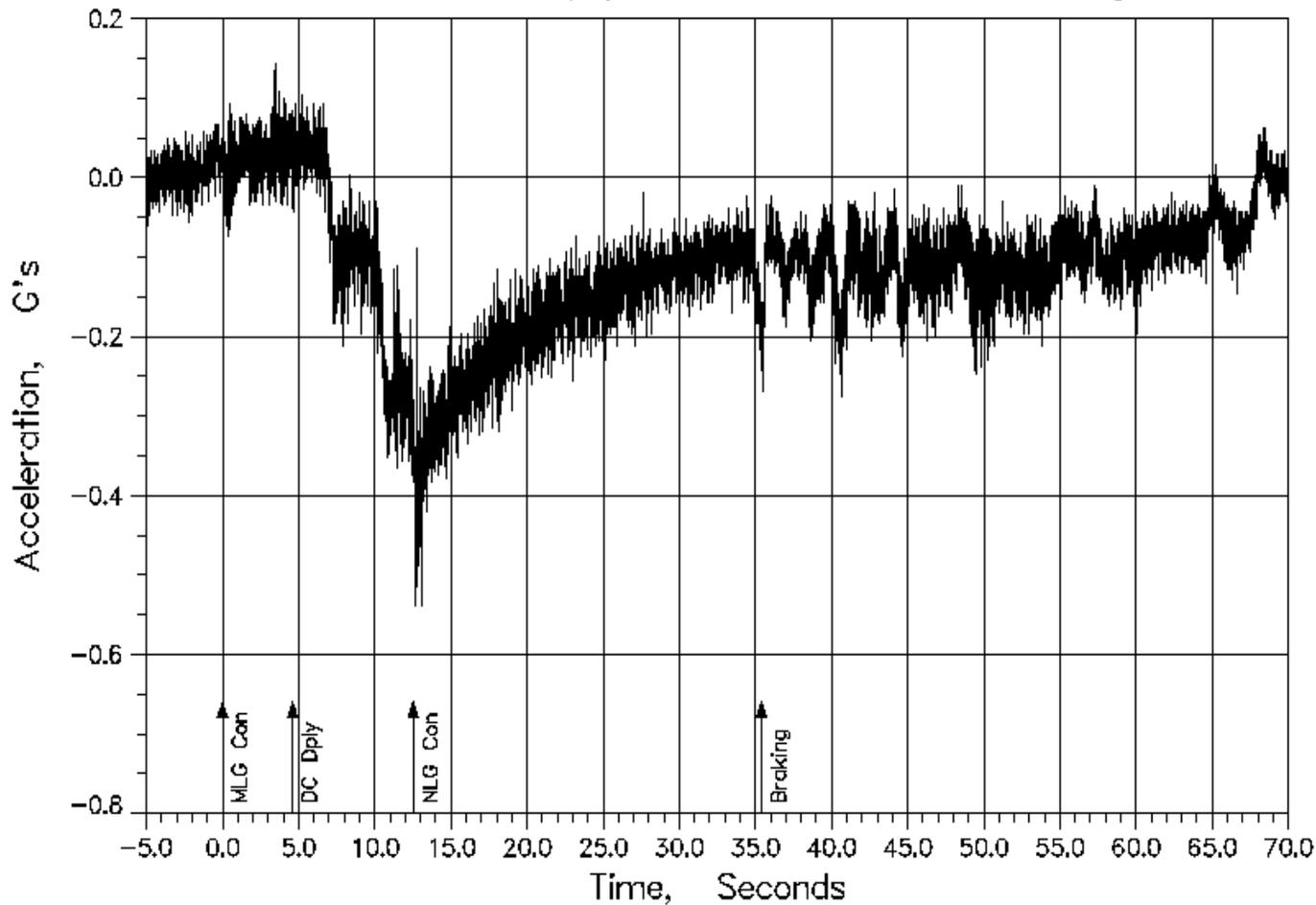


STS LANDING PROFILE

STS-92 Payload Bay Landing Time History

V34A9469A, DOF: X, Location: x=878.0, y=-102.0, z=407.0

MLG Contact=0.00s, DC Deploy=4.60s, NLG Contact=12.52s, Braking=35.40s

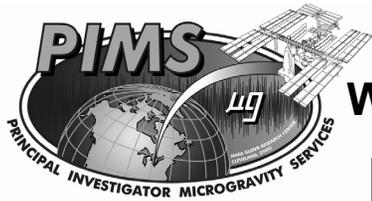


STS = Discovery

MLG_{cont} = main landing gear ground contact

DC_{deploy} = drag Chute deployment

NLG_{cont} = nose landing gear ground contact



Working in a Reduced Gravity Environment : "A Primer"

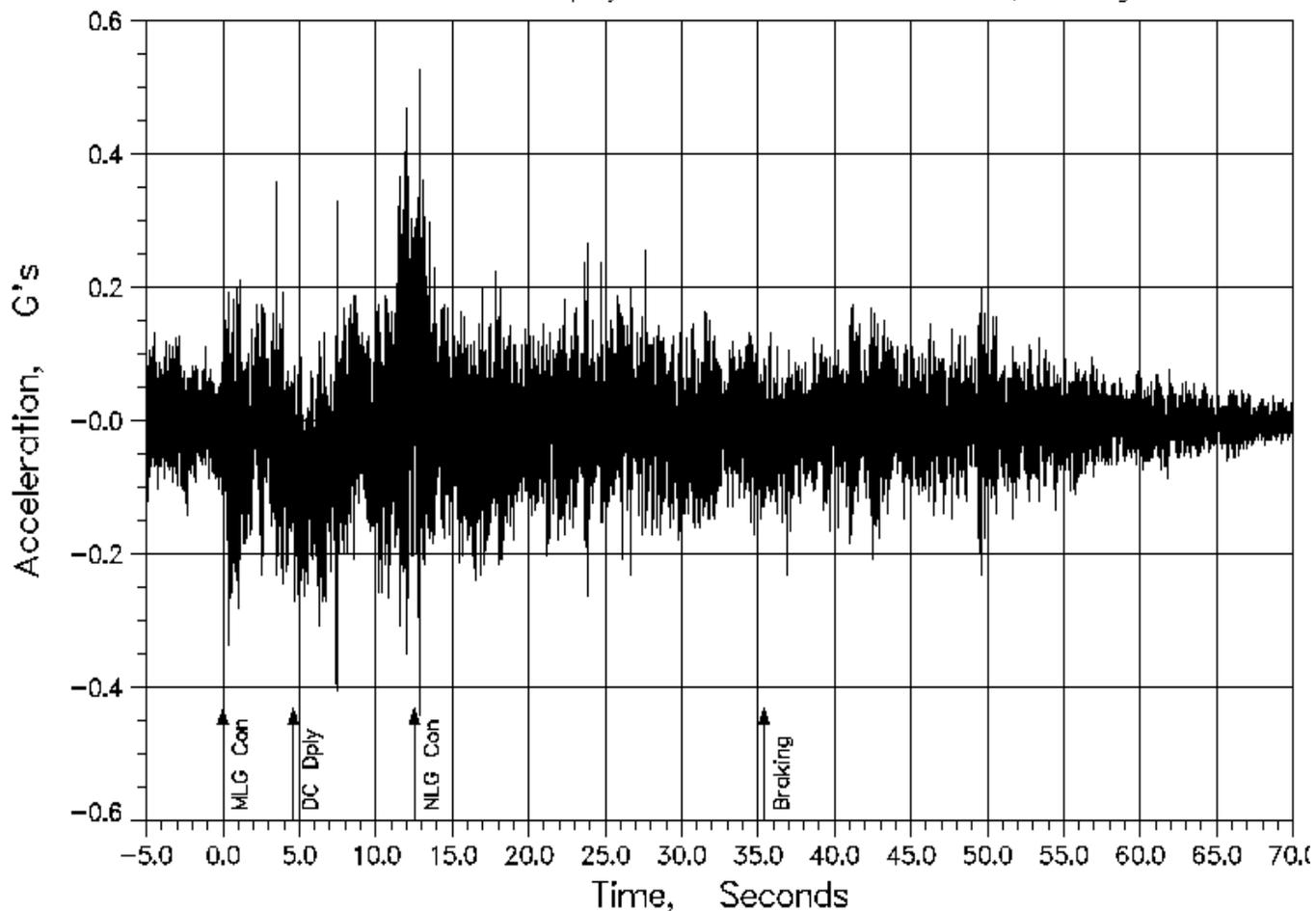


STS LANDING PROFILE

STS-92 Payload Bay Landing Time History

V34A9460A, DOF: Y, Location: x=701.0, y=-102.0, z=407.0

MLG Contact=0.00s, DC Deploy=4.60s, NLG Contact=12.52s, Braking=35.40s





Working in a Reduced Gravity Environment : "A Primer"

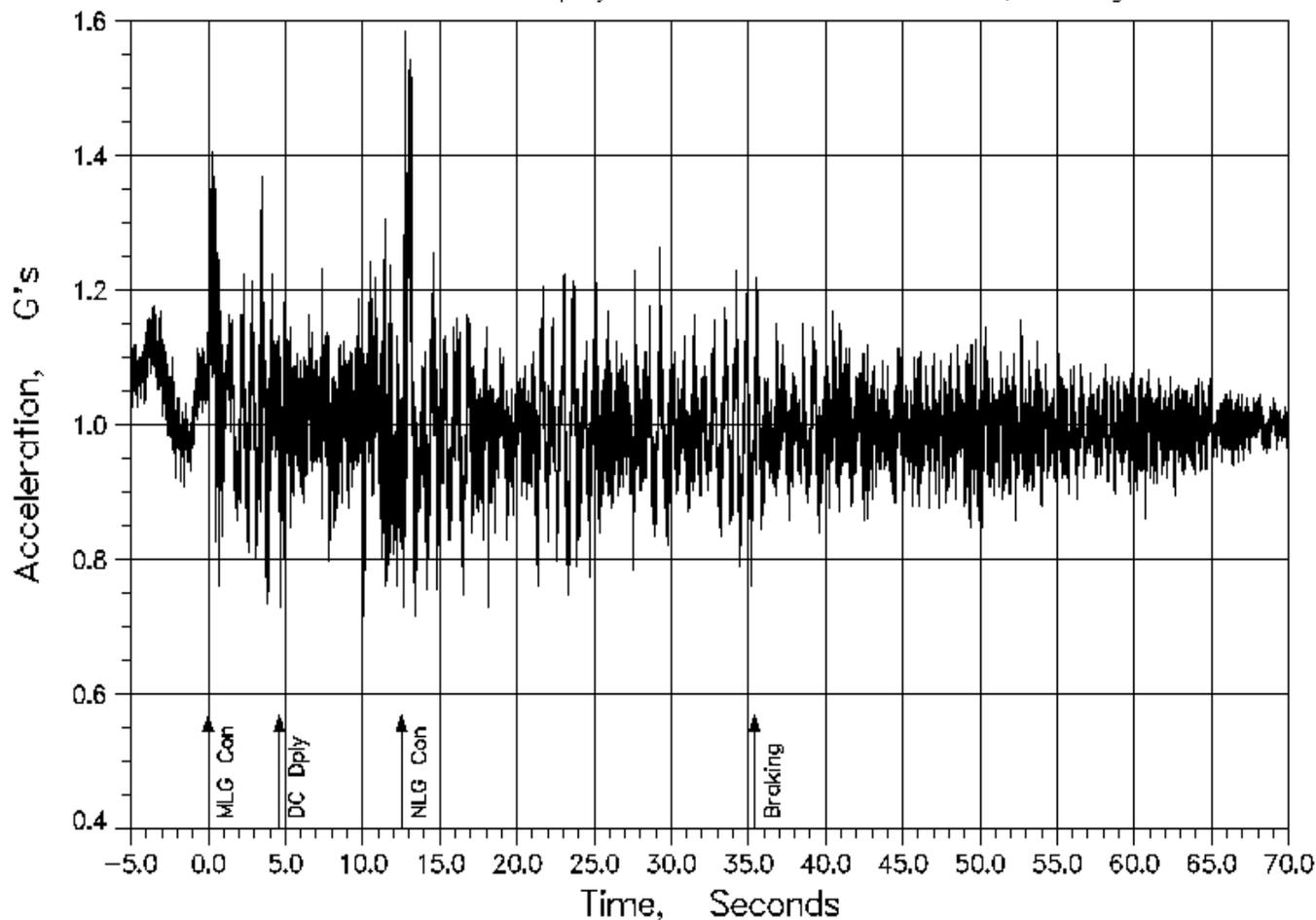


STS LANDING PROFILE

STS-92 Payload Bay Landing Time History

V34A9461A, DOF: Z, Location: x=701.0, y=-102.0, z=407.0

MLG Contact=0.00s, DC Deploy=4.60s, NLG Contact=12.52s, Braking=35.40s





Working in a Reduced Gravity Environment : "A Primer"

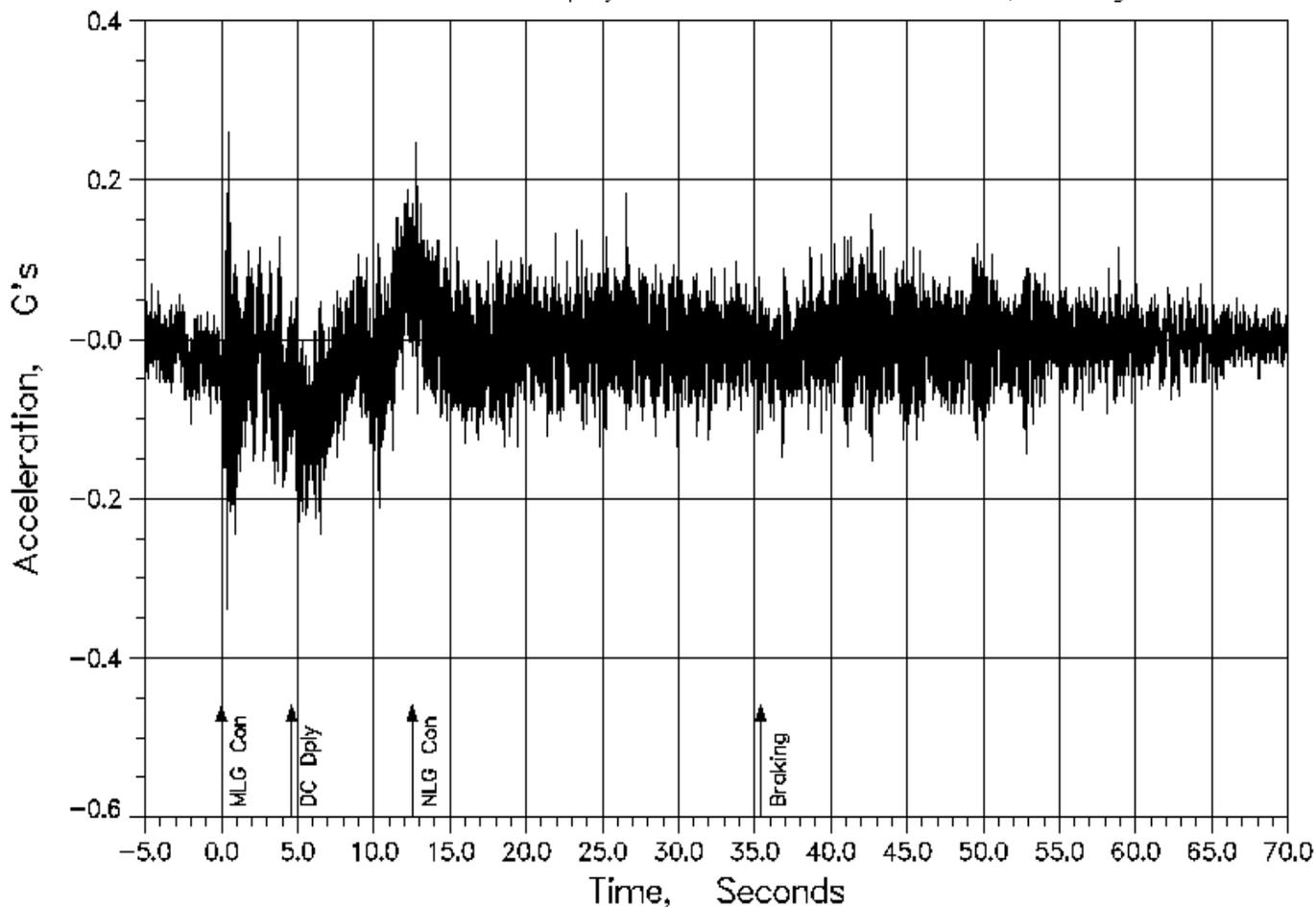


STS LANDING PROFILE

STS-92 Payload Bay Landing Time History

V34A9480A, DOF: Y, Location: x=919.0, y=-7.0, z=305.0

MLG Contact=0.00s, DC Deploy=4.60s, NLG Contact=12.52s, Braking=35.40s





Working in a Reduced Gravity Environment : “A Primer”



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